

The Protagonists

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The Atacama Large Millimeter/submillimeter Array

- Started Early Science operations in September 2011
 - 16 antennas
 - v = 84-720 GHz
 - Data rates: < 15 MB/s
- On completion:
 - 66 antennas
 - Data rates: 6 MB/s (avg) 60 MB/s (peak)
 - Majority of data consists of visibilities, but small fraction (0.1 MB/s) are Metadata and Auxiliary data.



Metadata & Auxiliary Data

- Metadata: describes the observation, visibilities, calibration and auxiliary data
 - eg. project number, spectral setup, location of data in the archive
- Auxiliary data: data required to process and analyse the primary data
 - eg. pointing, square-law detector, water-vapour radiometer data.
- Stored in the Archive, as ALMA Science Data Model (ASDM)

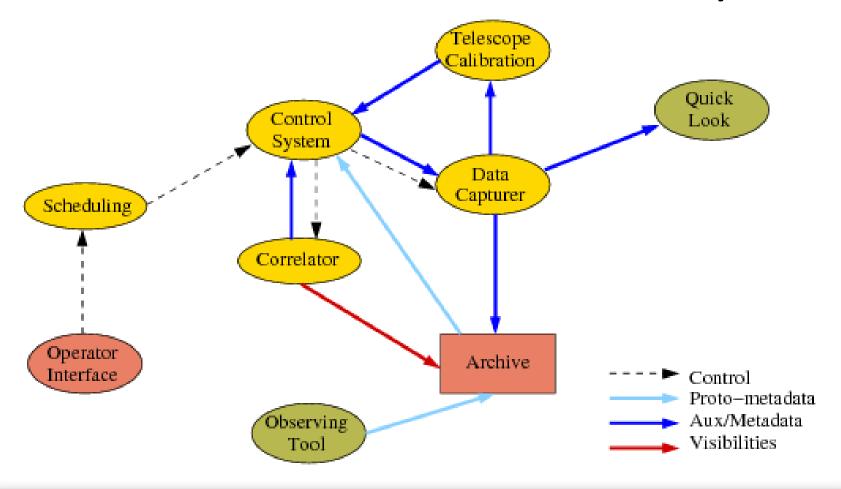


ALMA Science Data Model (SDM)

- Must contain all the information necessary for the astronomical processing of raw data from the telescope.
- Format in which raw science data will be archived and provided to astronomical observers.
- Supported by the Common Astronomy Software Applications (CASA) post-processing software
- Set of tables represented as XML documents (mostly)
- Data from correlators, square-law detectors, radiometers, etc. stored in binary blocks.



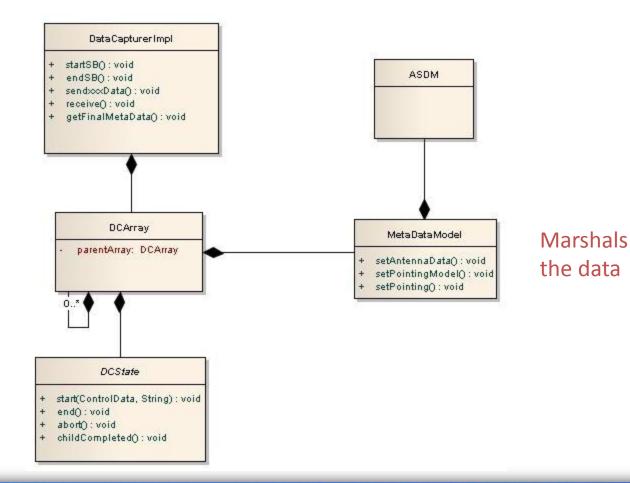
Data Flow in the ALMA software system





Capturing the Meta/Auxiliary Data

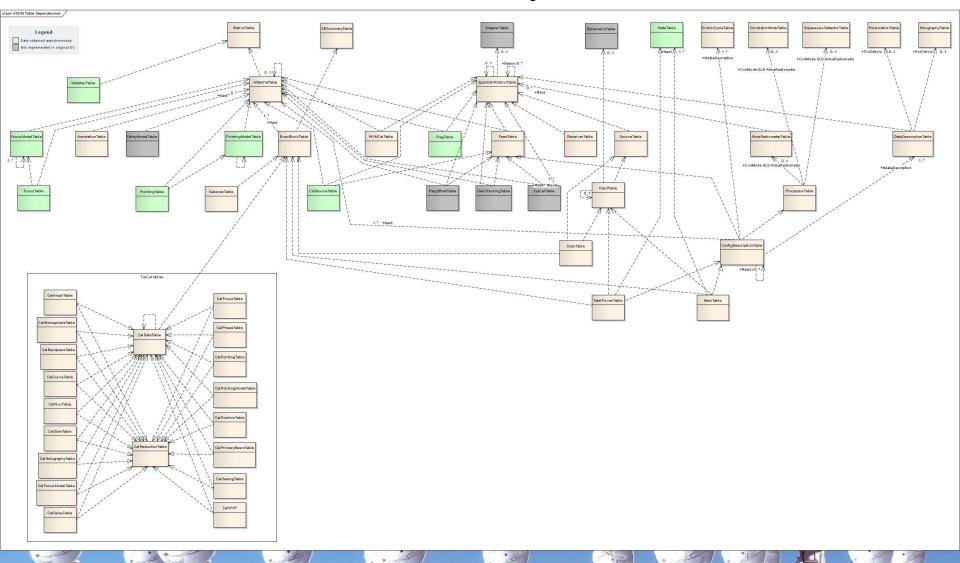
Interface to other subsystems



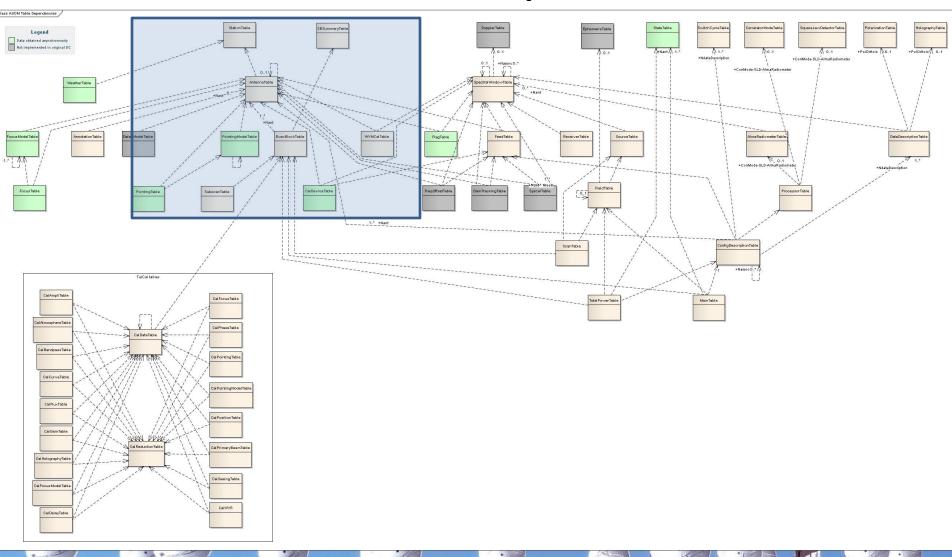
State machine



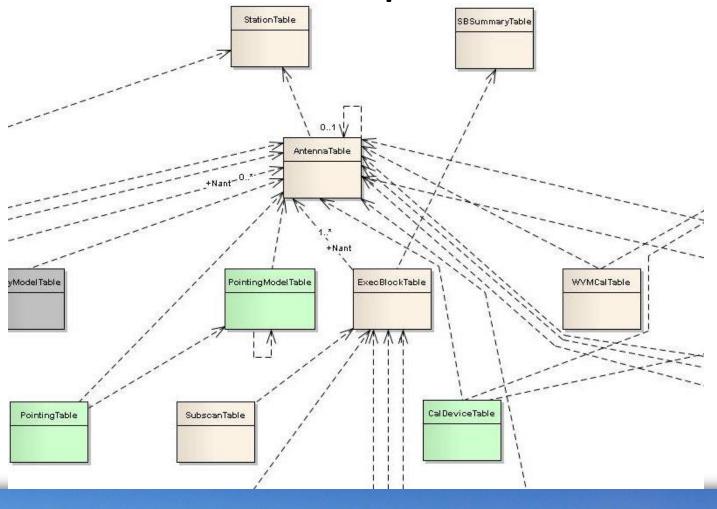
ASDM Table Dependencies



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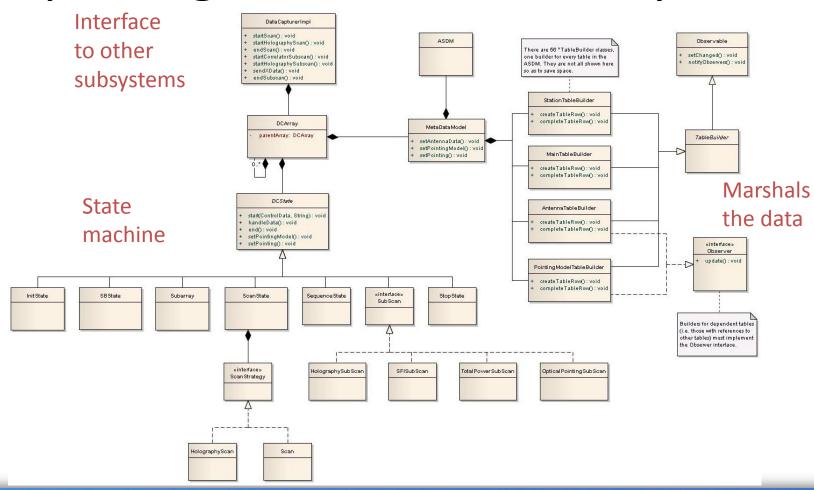








Capturing the Meta/Auxiliary Data





What have we gained?

- Maintainability: uses several specialised data or functionbased classes, rather than a very small number of classes to do everything.
- Consolidation: have been able to remove much code repetition (eg. for different observing modes).
- Extensibility: for new ASDM tables, new observing modes
- Better testing: specialisation of tasks results in a module which is easier to test (at the unit and end-to-end levels)





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