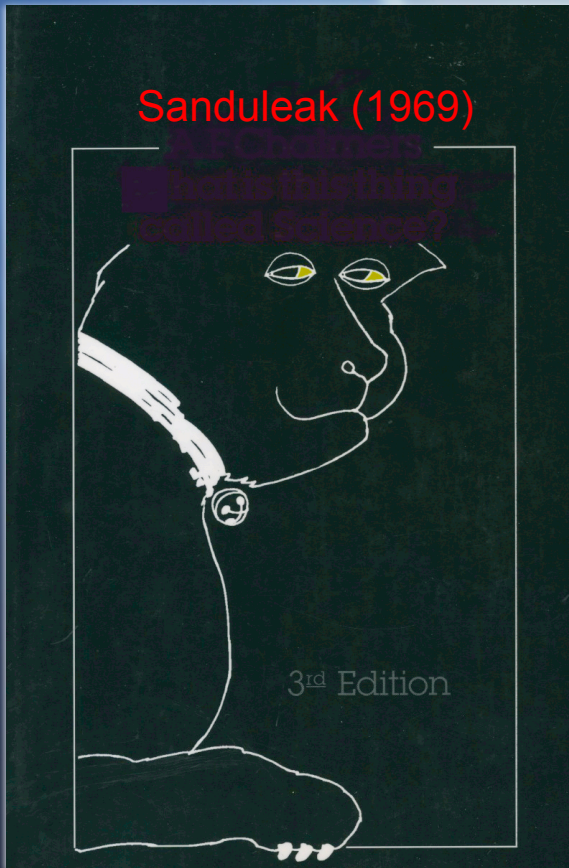
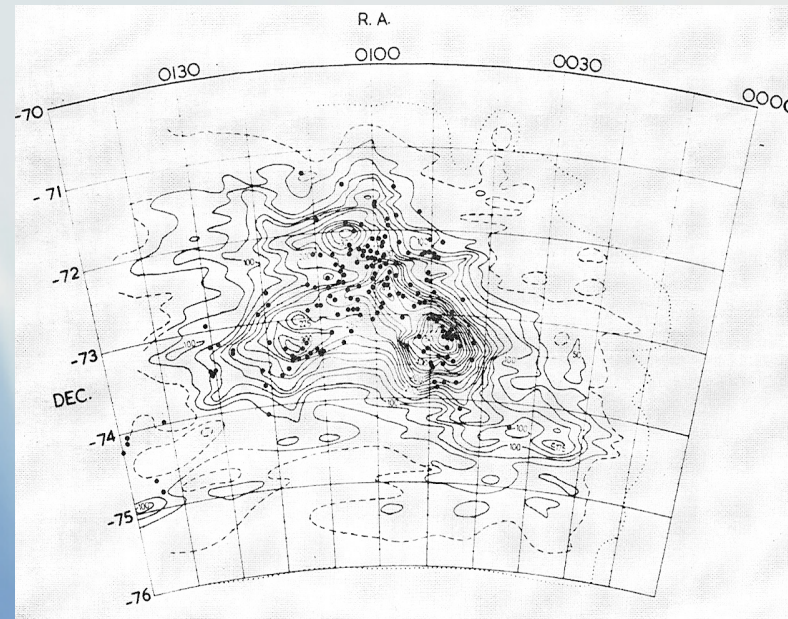
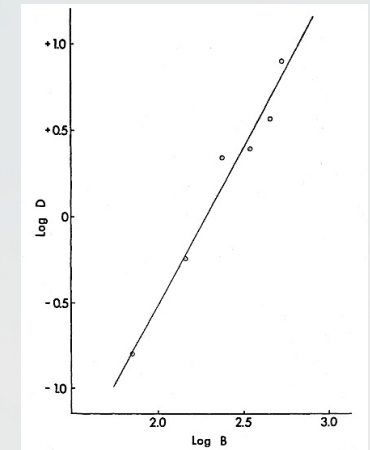


“The Rate of Star Formation”

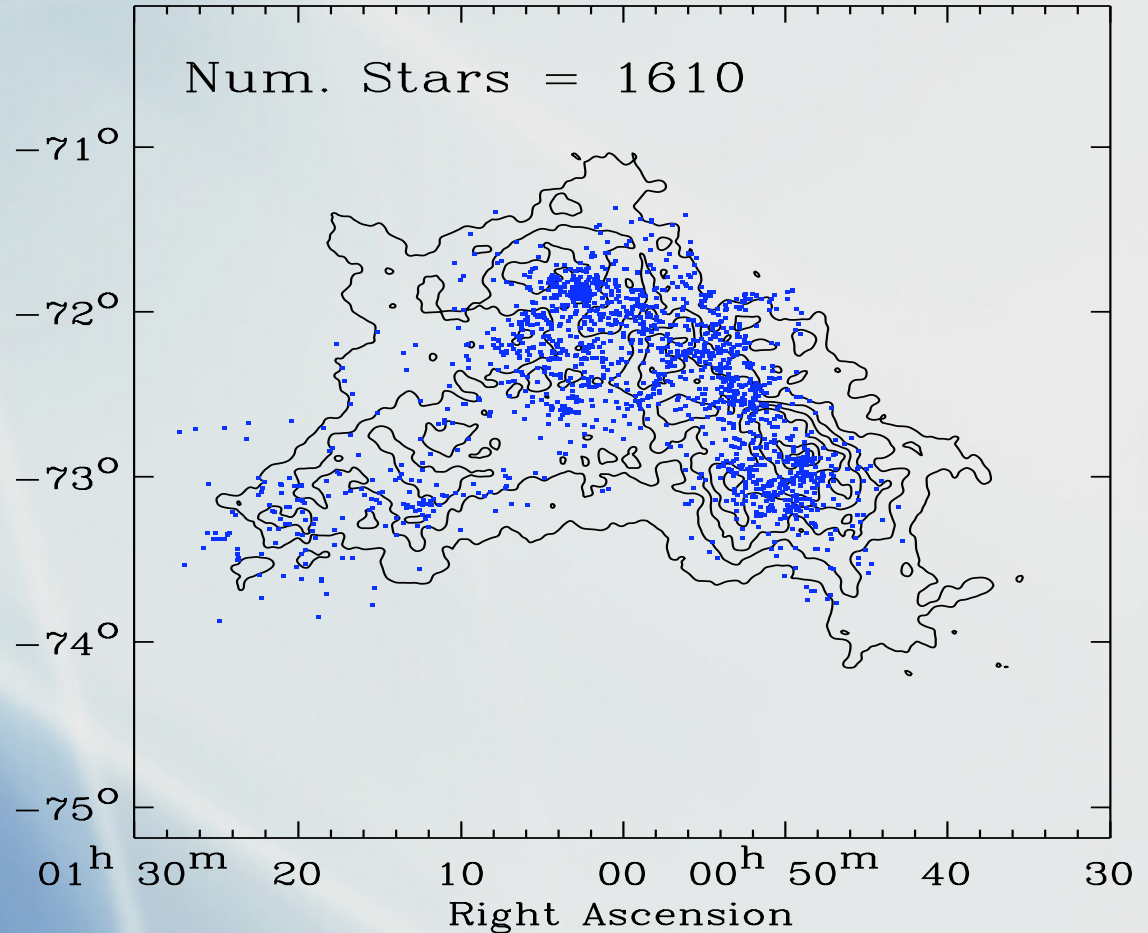
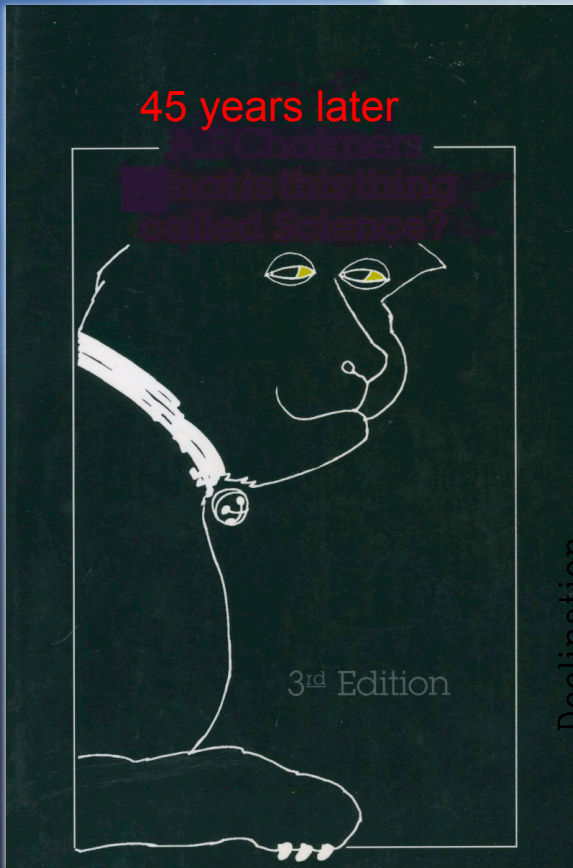
M. Schmidt vs. J. Einasto

This is what started it all.

A Neutral Hydrogen map and
170 Individually Resolved
OB Stars in the SMC
Sanduleak (1969)

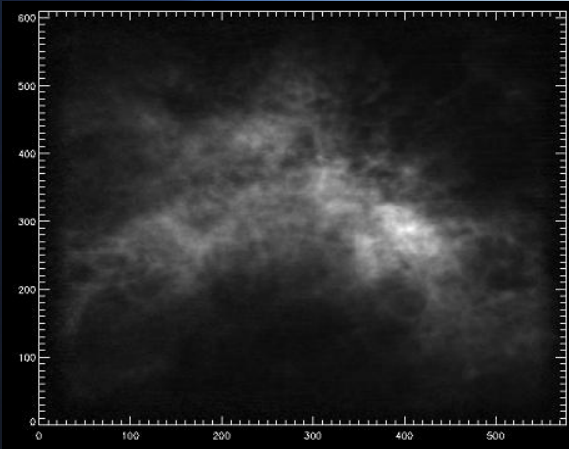


Back to the SMC (2011)

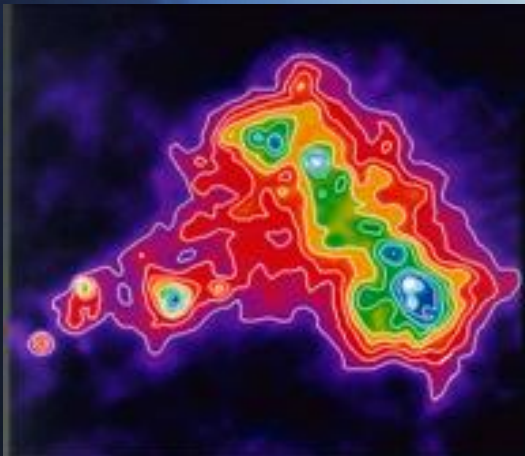
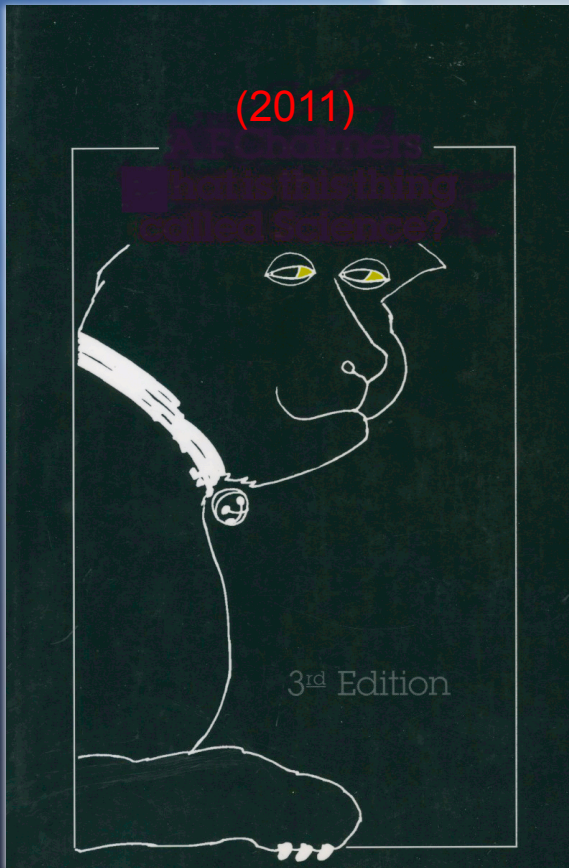
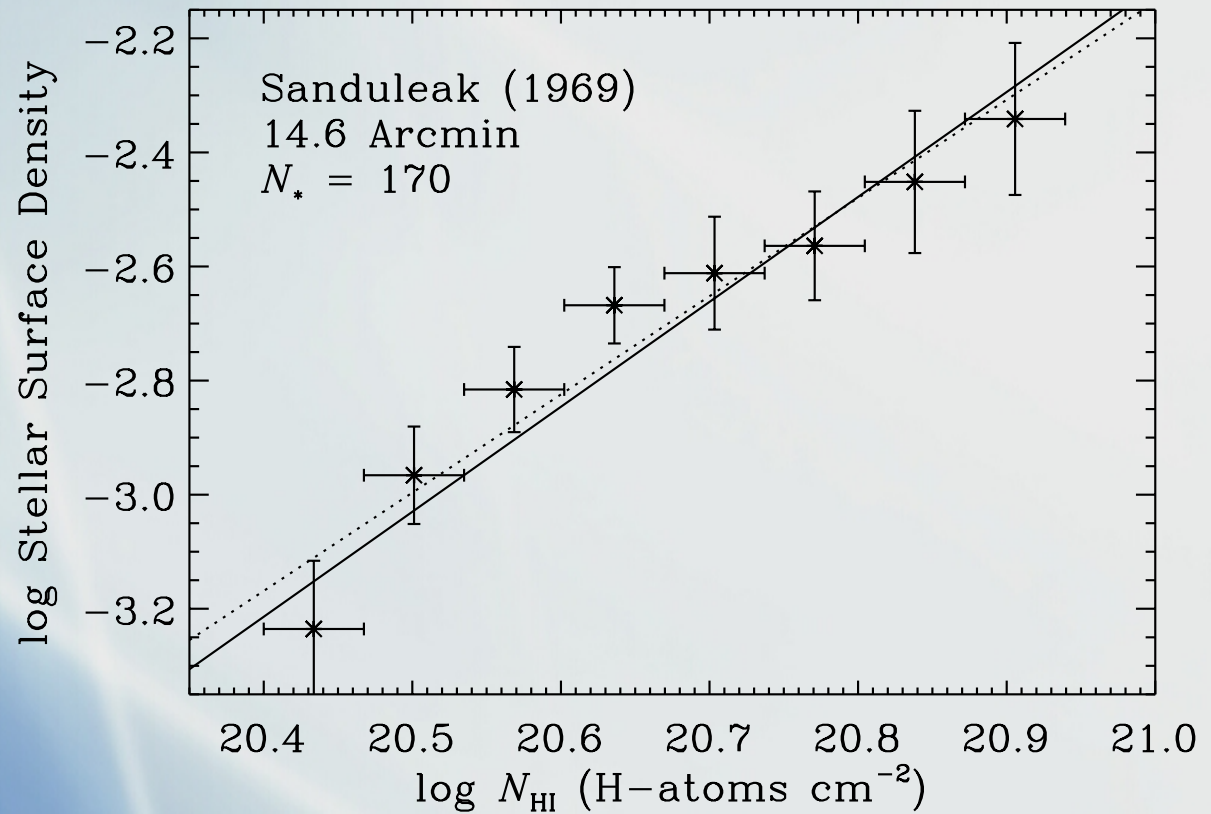


Center: R.A. 01 00 56.94 Dec -72 53 25.5

UBVR (85K) Catalog: Massey, P. 2002, ApJS, 141, 81
HI Map: Kim et al. 1998, ApJ, 503, 674



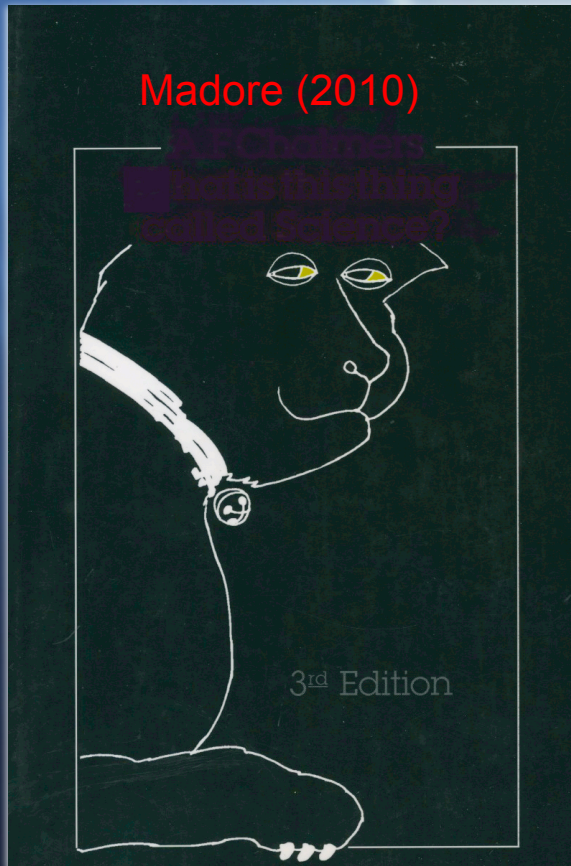
Star Formation traced by OB Supergiants



Cloud-Collapse and Star-Formation

Stagnation Timescales

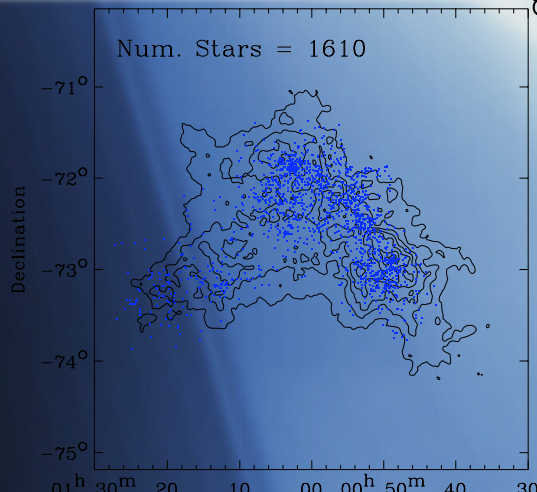
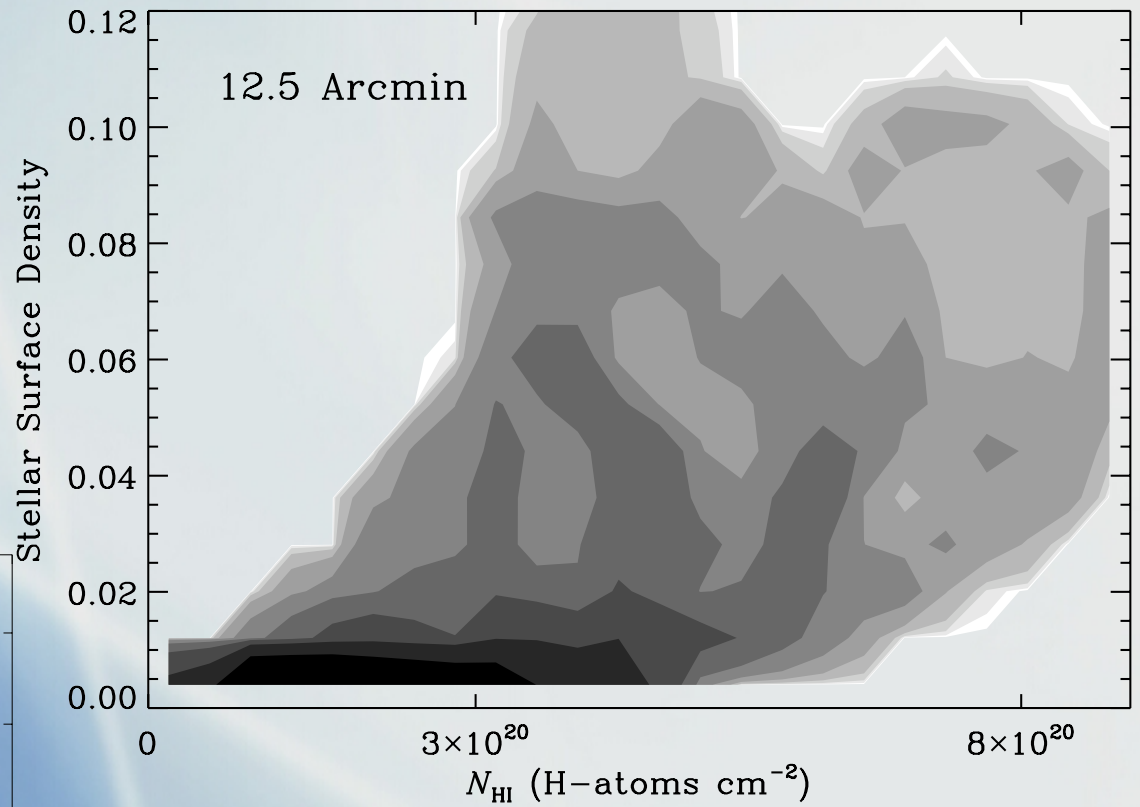
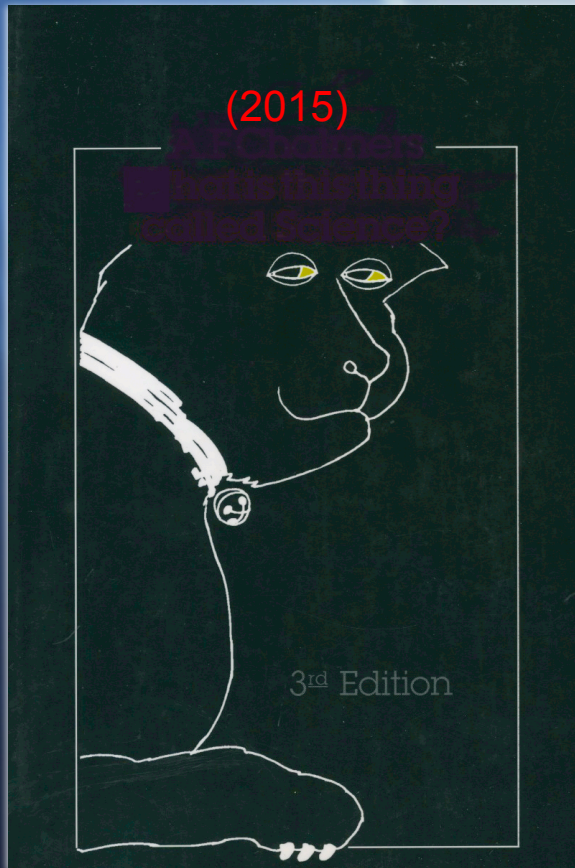
(Madore, B.F., 2010, ApJL, 716, 131)



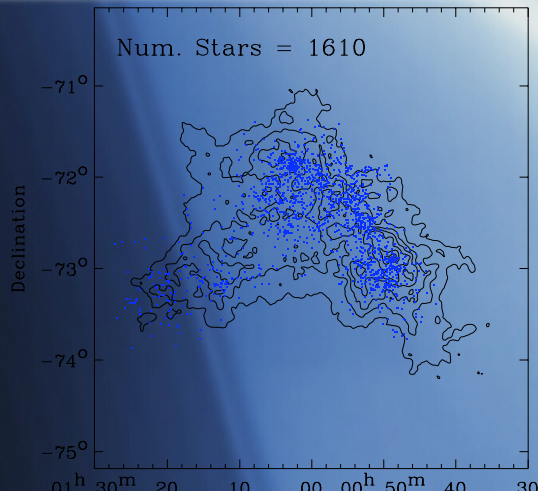
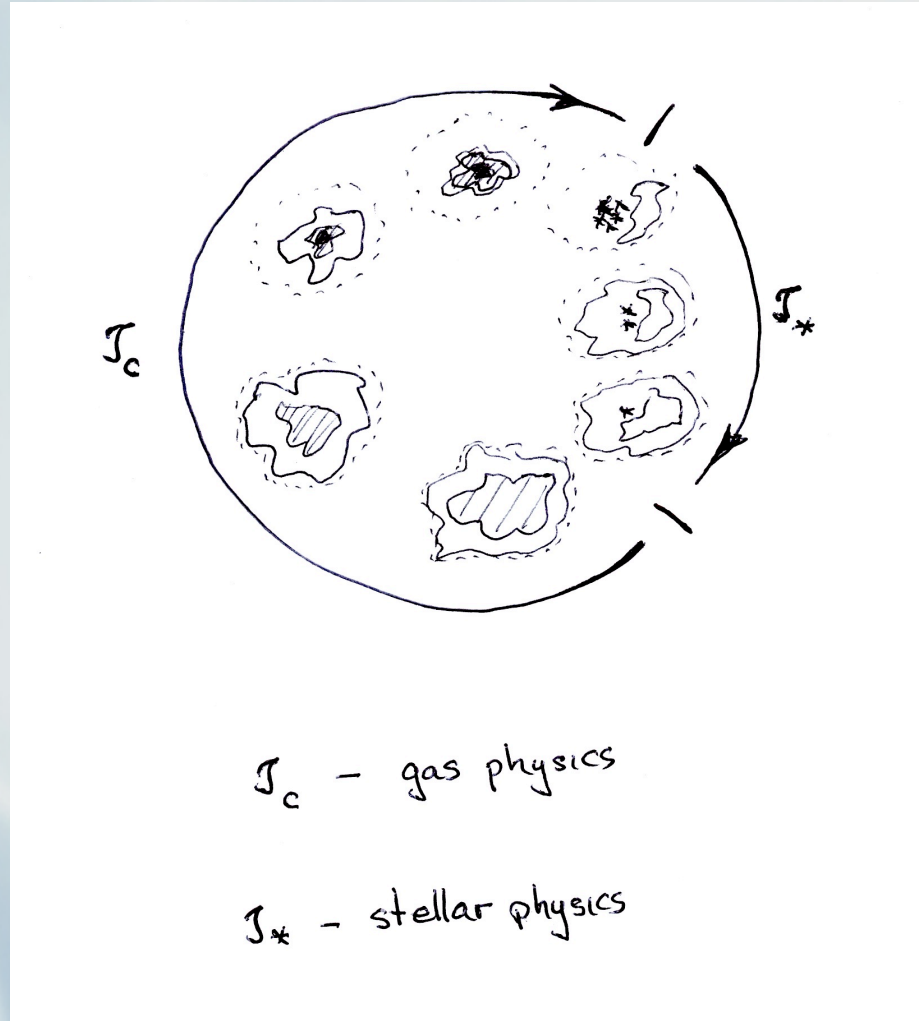
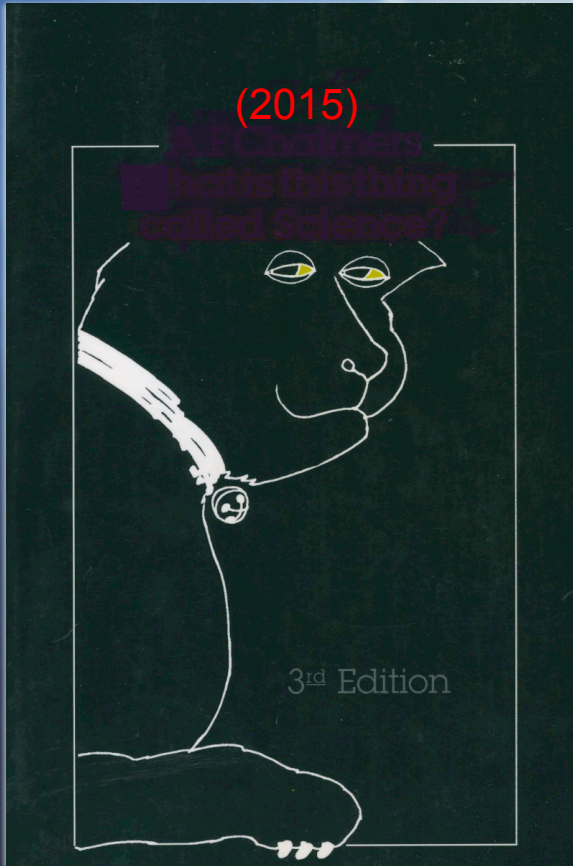
$$\begin{aligned} f_{SFR} &= \frac{M_*}{\tau} = \frac{\epsilon V \rho_g}{(\tau_* + \tau_c)} \\ &= \frac{A \rho_g}{(\tau_* + B(\rho_g/\rho_c)^{-n})} \\ &= \frac{A \rho_g / \tau_*}{(1 + (B \rho_c^n / \tau_*) \rho_g^{-n})} \\ &= \frac{\alpha \rho_g}{(1 + \beta \rho_g^{-n})} \quad (1) \end{aligned}$$

where $A = \epsilon V$, $\alpha = A/\tau_*$, $\beta = (B \rho_c^n \tau_*)$ and it is assumed that $\tau_c = B(\rho_g/\rho_c)^{-n}$.

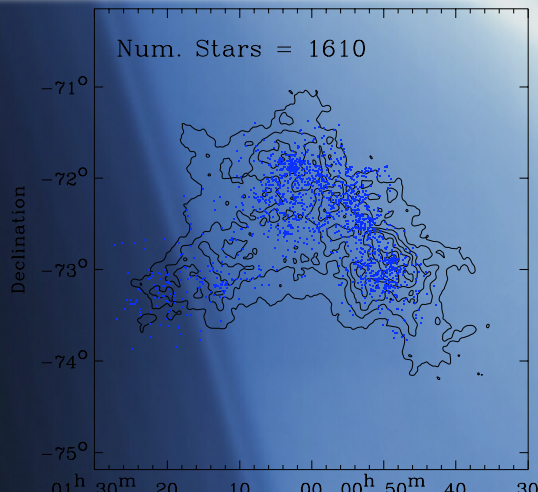
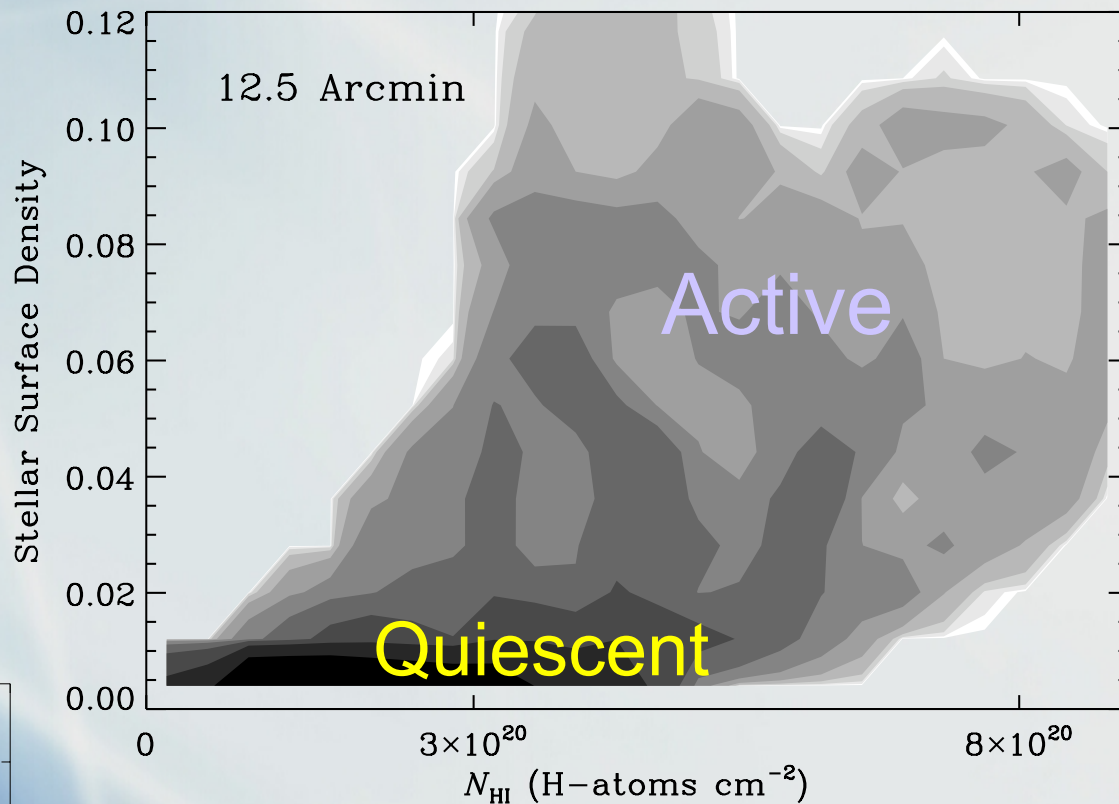
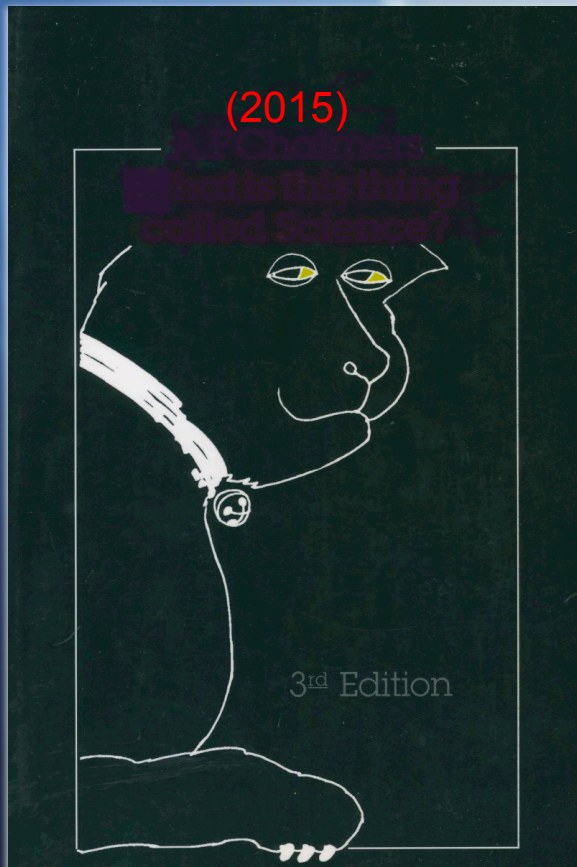
The Hess Diagram for Star Formation

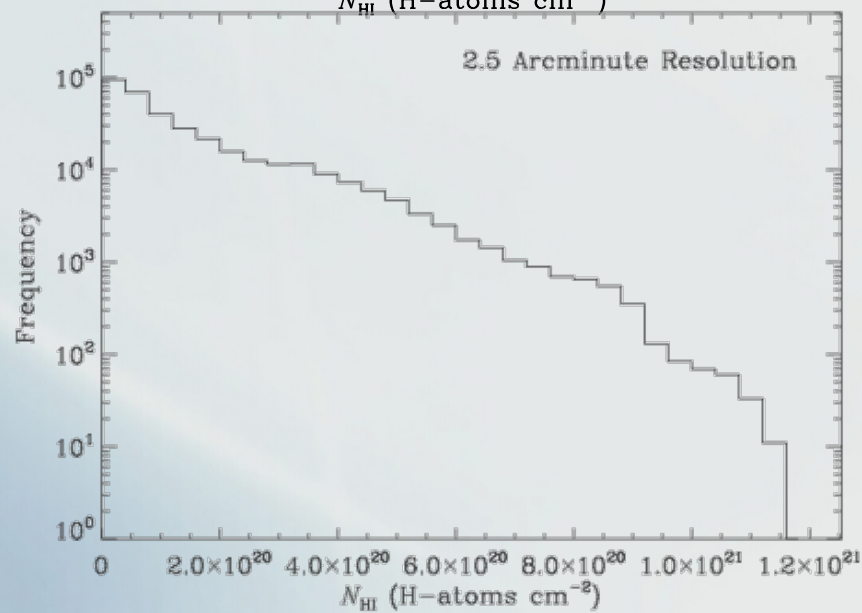
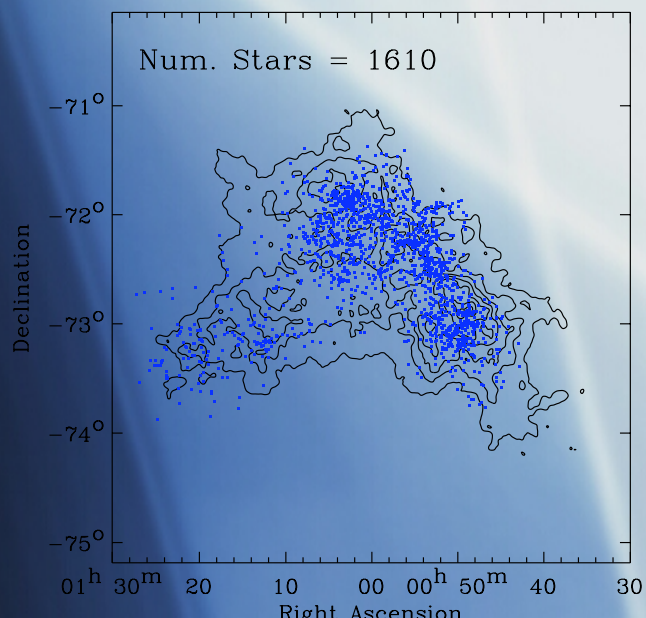
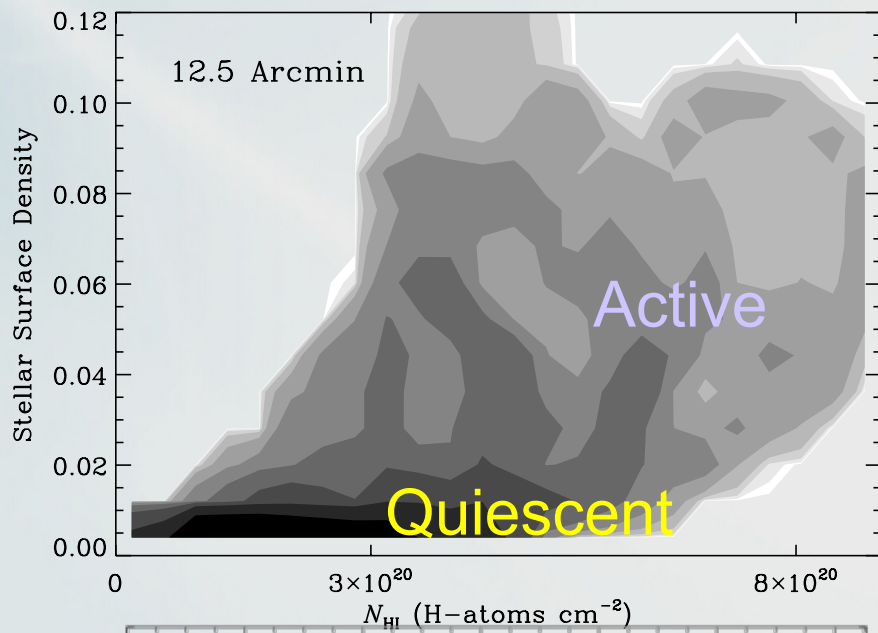
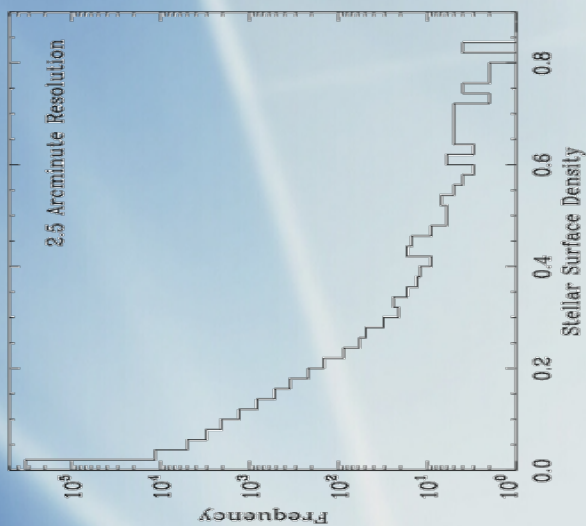


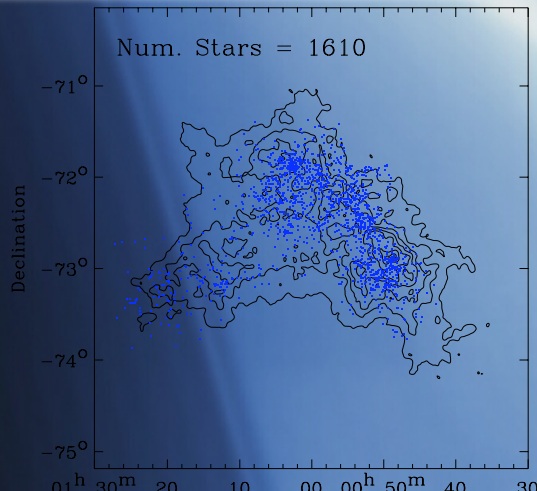
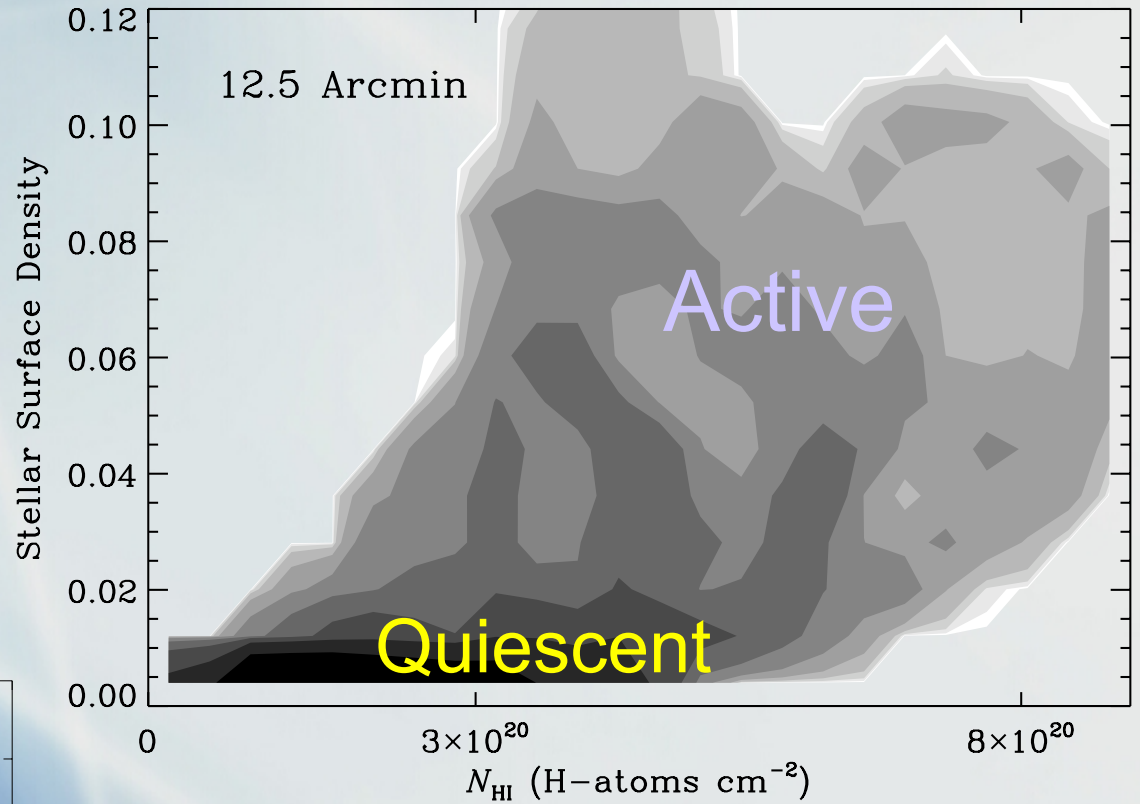
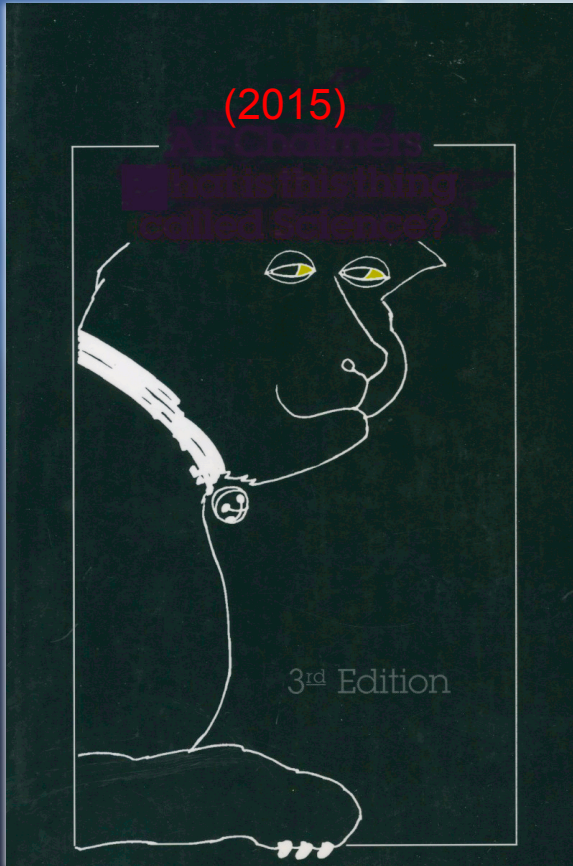
The Cycle of Star Formation



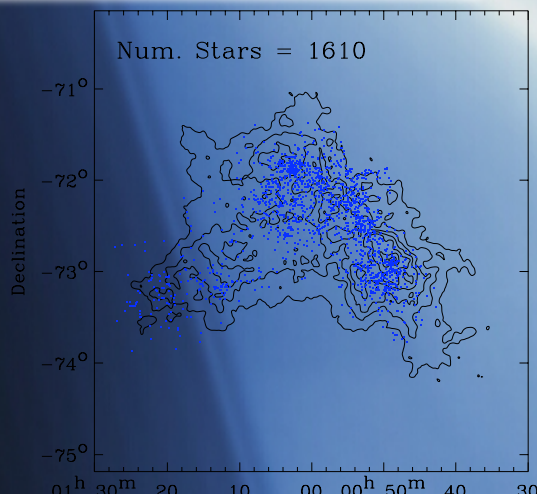
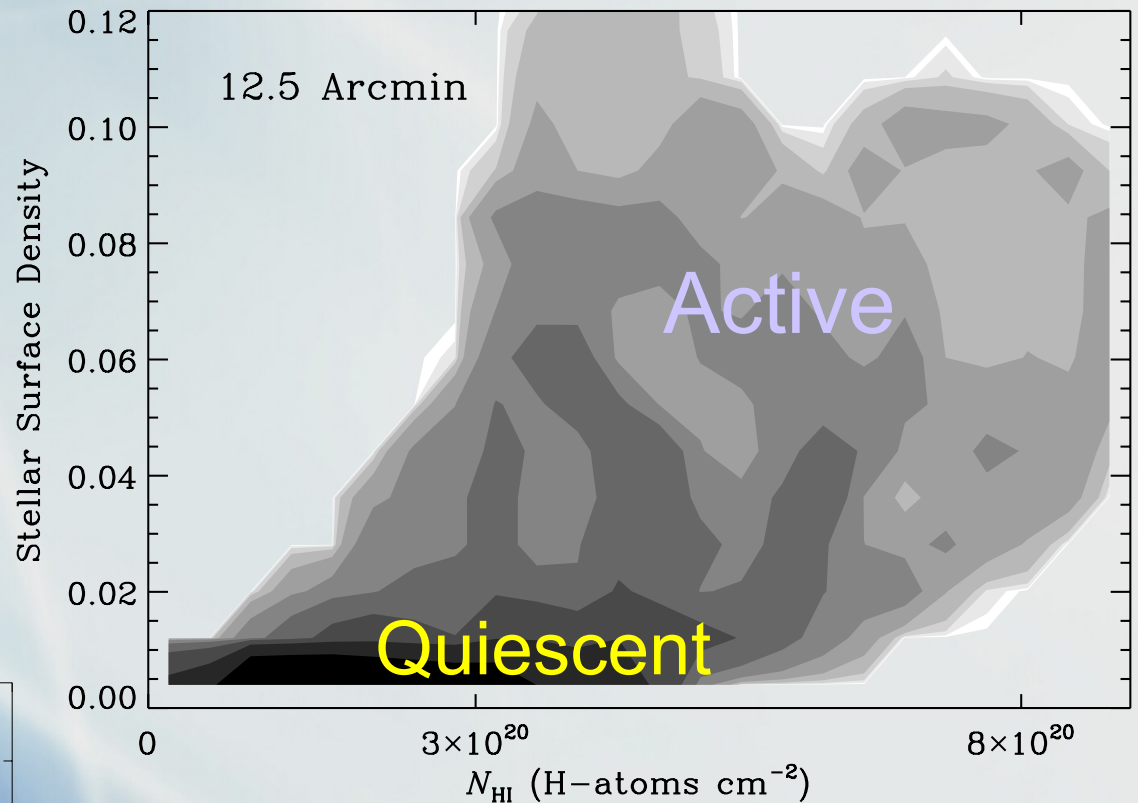
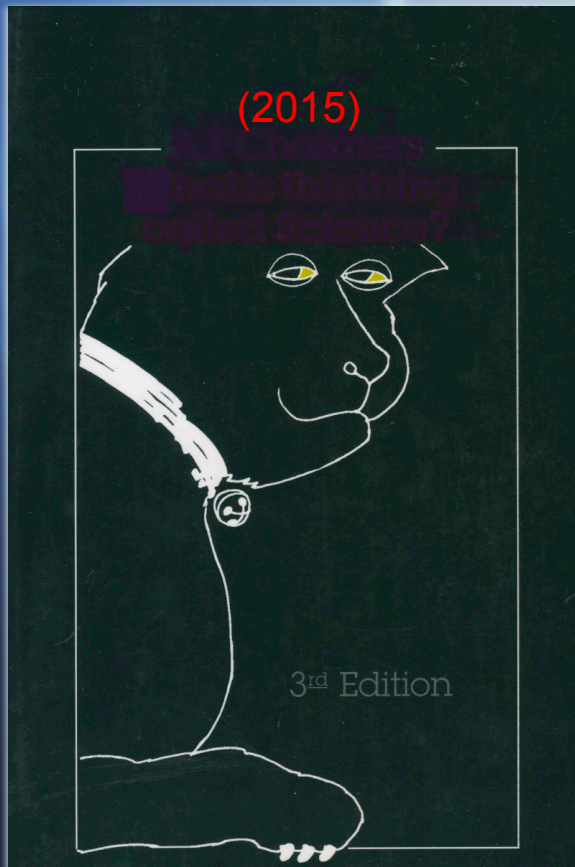
Rate-Limiting States







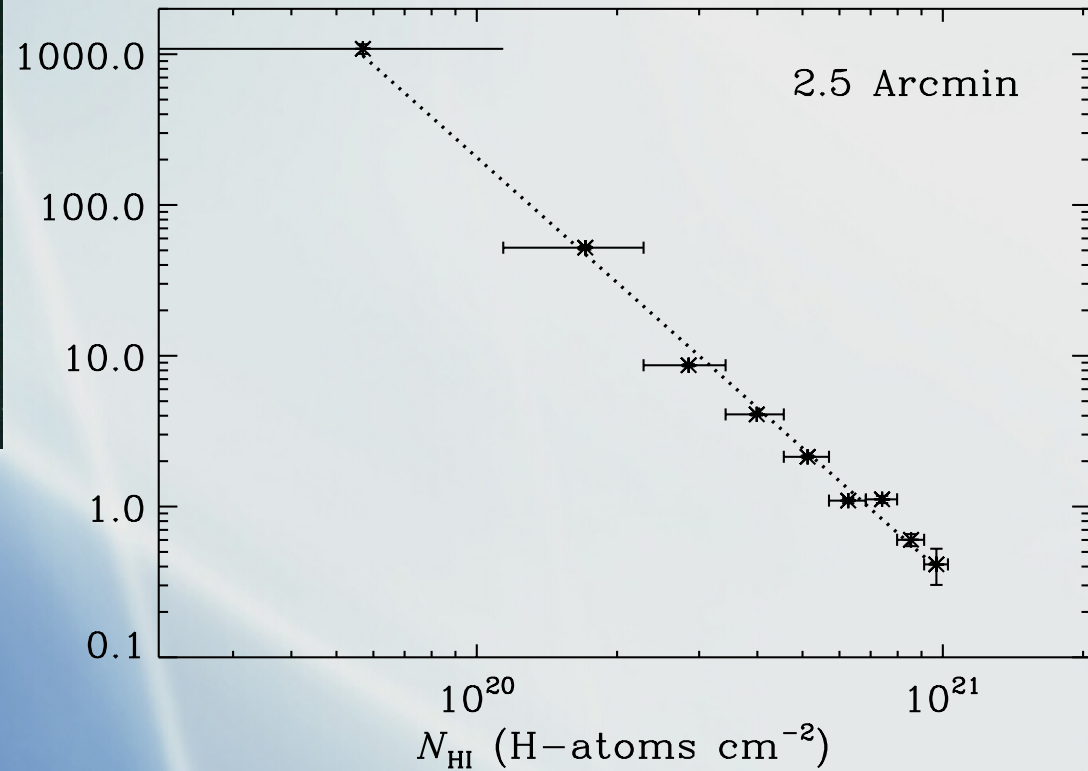
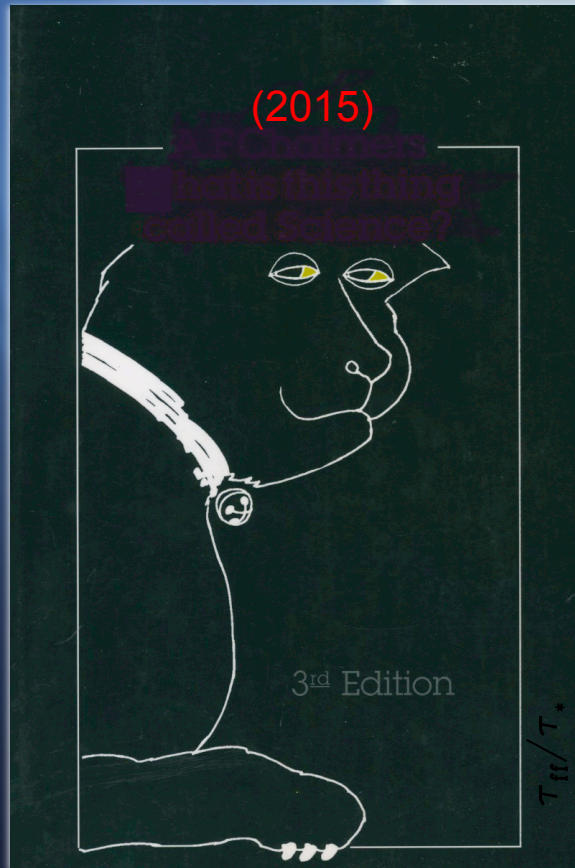
Areal frequencies → Relative timescales

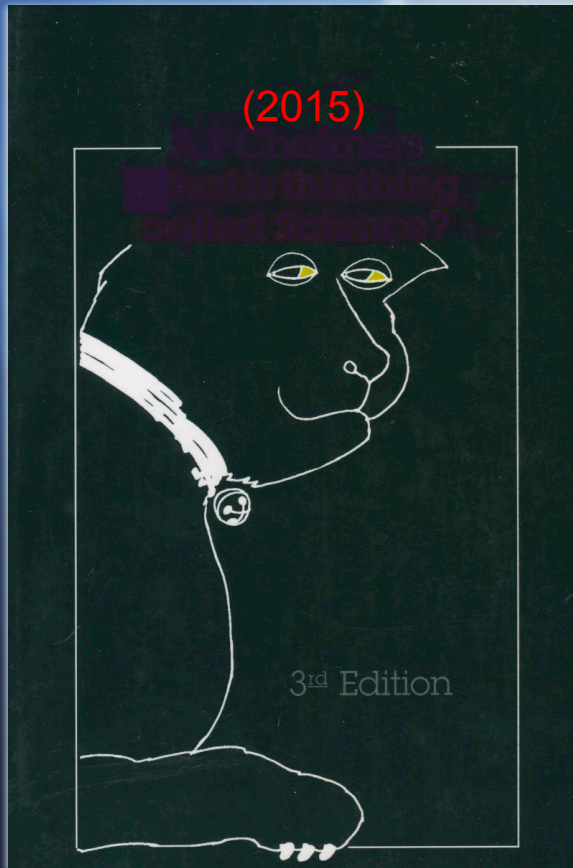


Areal frequencies \rightarrow Relative timescales

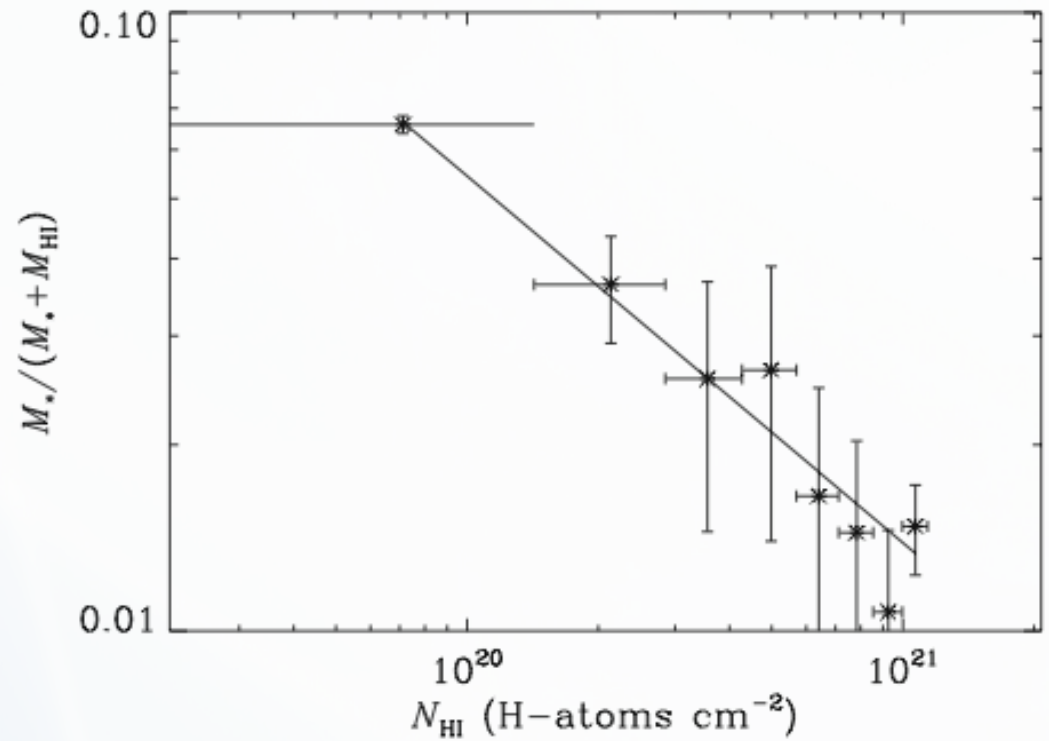
$$N(\text{quiescent})/N(\text{active}) = t(\text{cloud})/t(\text{stellar})$$

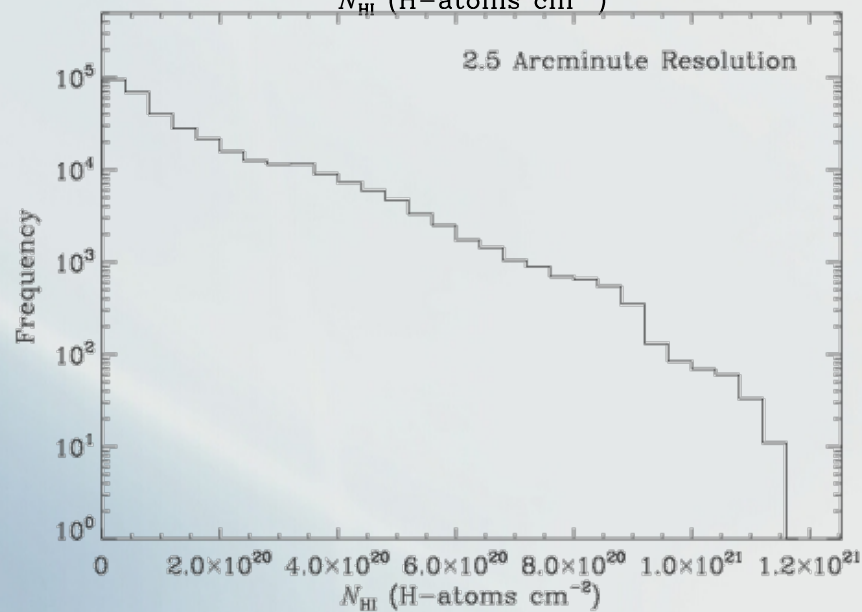
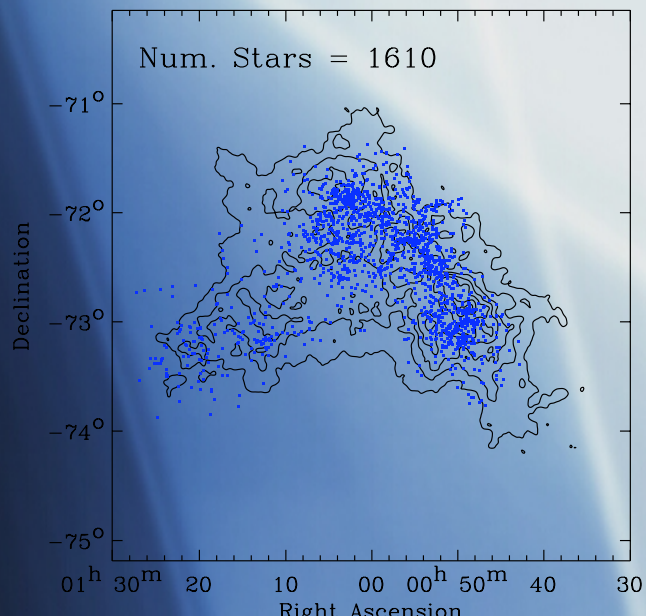
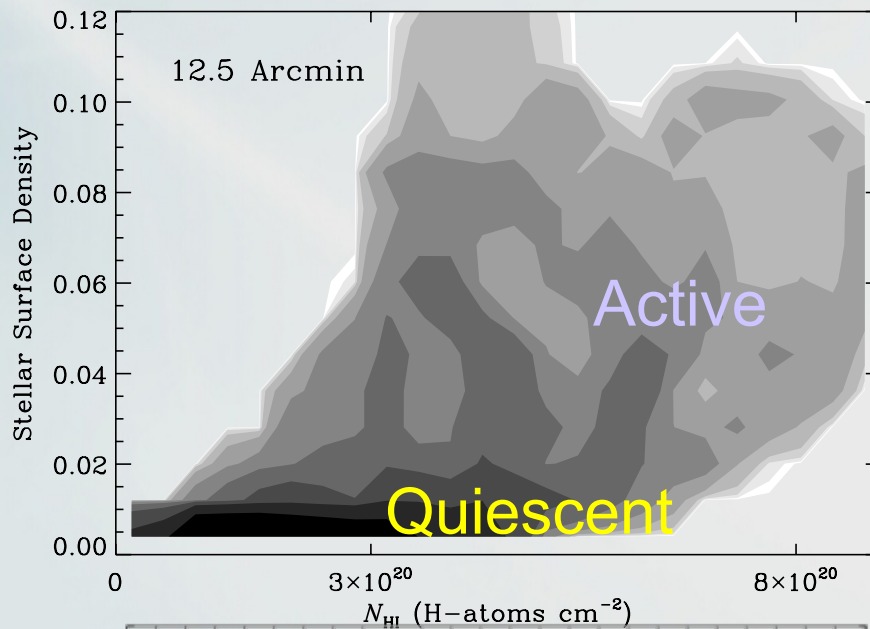
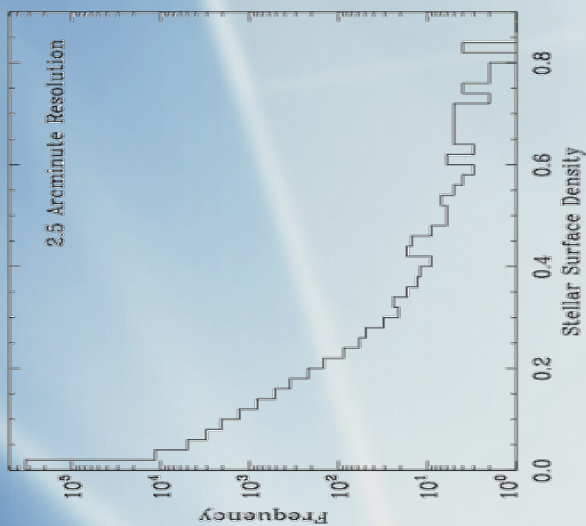
Timescales for Star Formation





Efficiency of Star Formation





“I really don't get it”

