

Mass determination of young directly imaged planets and brown dwarfs

Tobias Schmidt

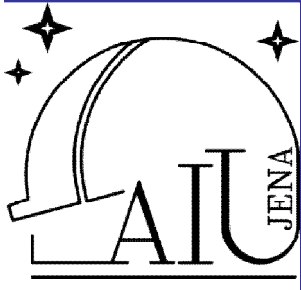
with Ralph Neuhäuser (Jena, Germany)

Andreas Seifahrt (Davis, USA)

Sören Witte, Peter Hauschildt (Hamburg, Germany)

Christiane Helling (St. Andrews, UK)

M. Mugrauer, K. Schreyer, S. Fiedler, Ch. Ginski, Ch. Adam (Jena)



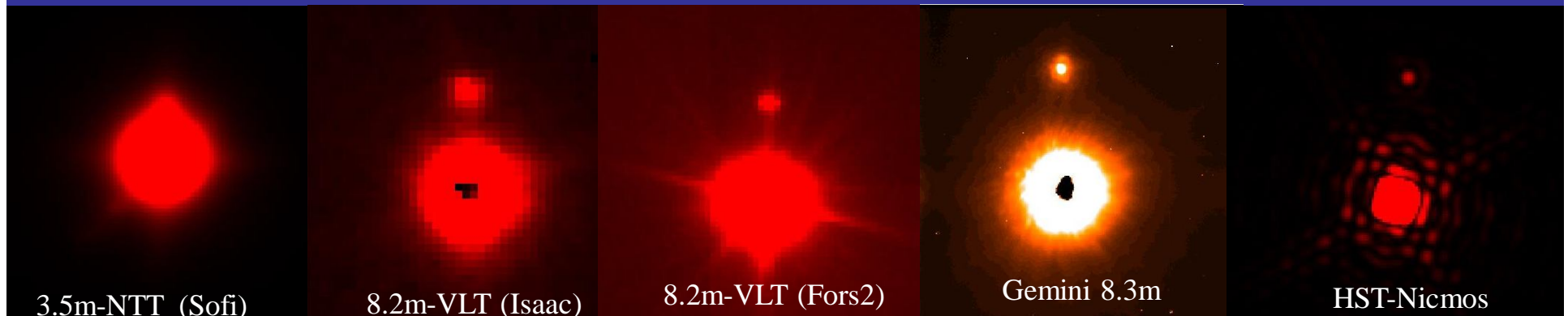
seit 1558

Astrophysical Institute and University Observatory

www.exoplanet.de

www.astro.uni-jena.de

Friedrich-Schiller-University Jena, Germany



3.5m-NTT (Sofi)

8.2m-VLT (Isaac)

8.2m-VLT (Fors2)

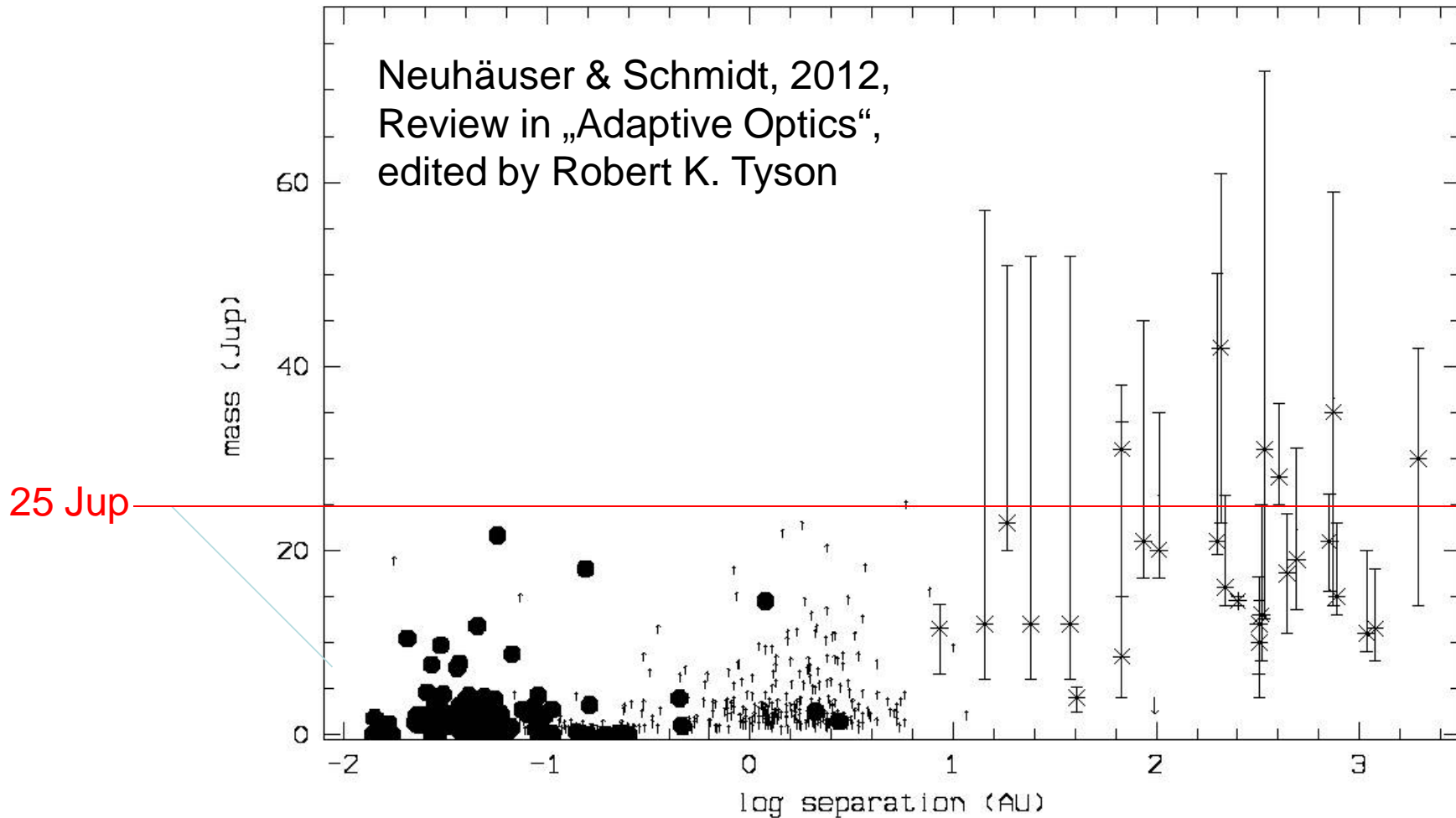
Gemini 8.3m

HST-Nicmos

methods now meet at ~ 10 AU

11 objects not yet on exoplanet.eu:

GG Tau Bb, TWA 5 B, GJ 417 B & C, GSC 08047 B/b, LP 261-75 B/b,
HD 203030 B/b, Wolf 940 B/b, G196-3 B/b, PZ Tel B/b, HR 7329 B/b,



RV planets at 9-12 AU: nu Oph c, HIP 5158 b, HIP 70849 b, 47 Uma d

Imaged planets at 9-18 AU: β Pic b, HR 8799 e, PZ Tel b

CT Cha b and Drift-Phoenix:

$T = 2600 \text{ K} \pm 250 \text{ K}$

$A_V = 5.2 \pm 0.8 \text{ mag}$

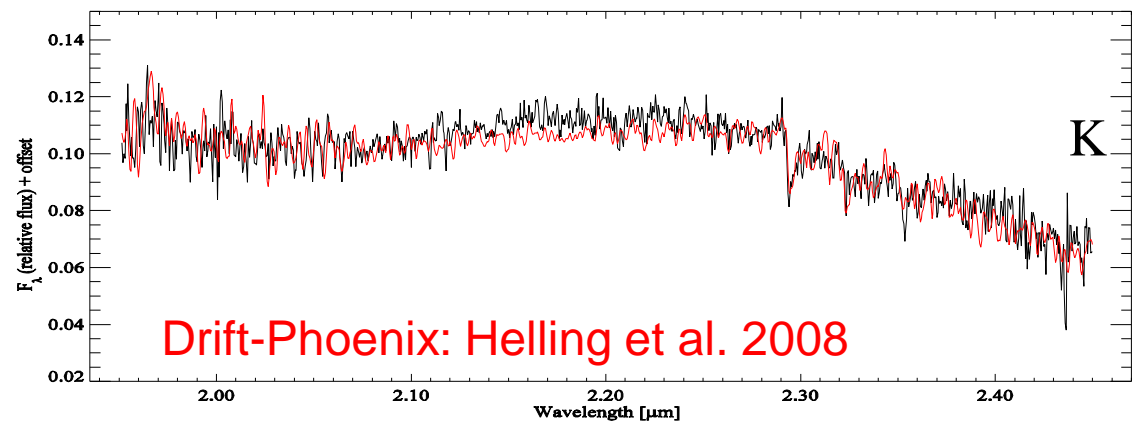
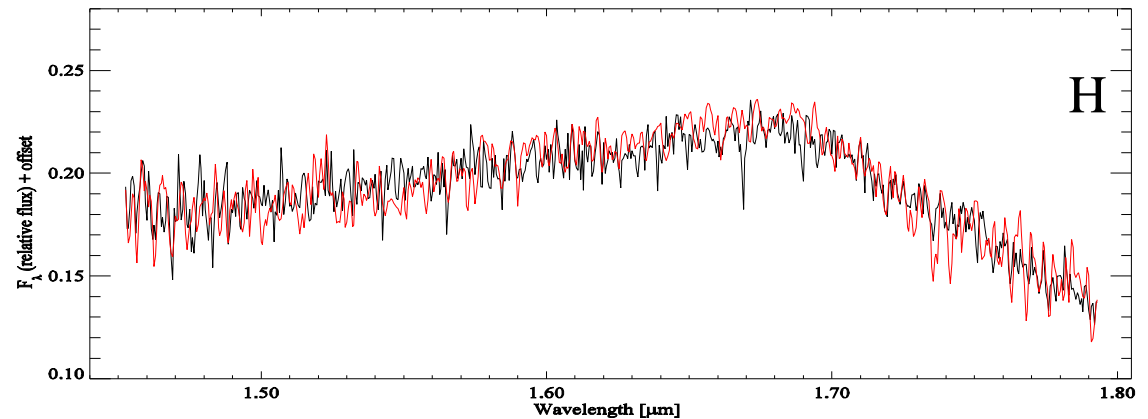
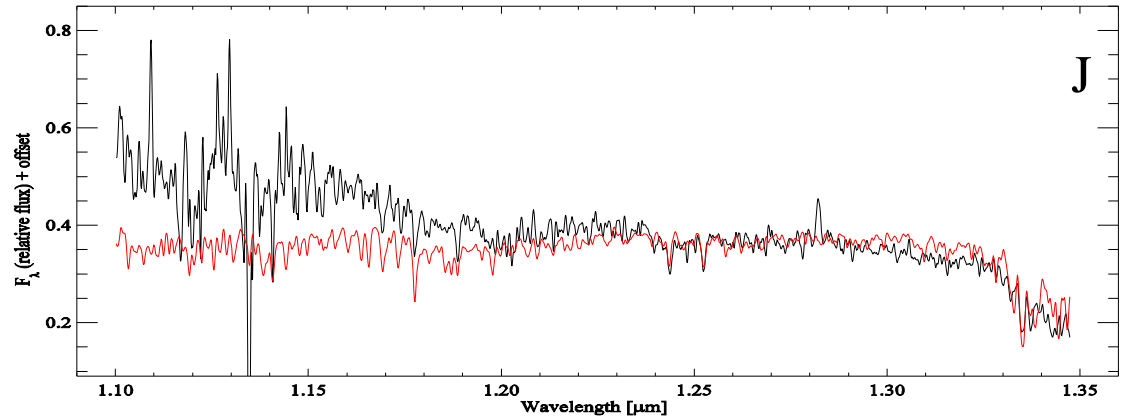
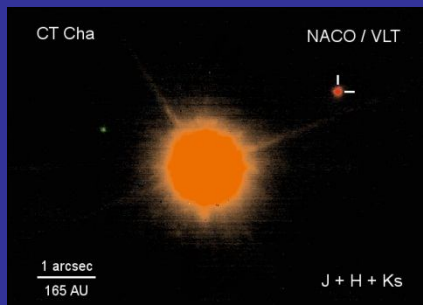
→ Mag, A_V and distance
give luminosity L
($\log(L/L_{\text{sun}}) = -2.68 \pm 0.21$)

→ L and T give radius
($\sim 2.2 \pm 0.7 R_{\text{Jup}}$)

→ T , L , age give mass:

$\sim 17 \pm 6 M_{\text{Jup}}$

but model-dependant
(planet or BD ?)



Schmidt, Neuhauser, Seifahrt et al. 2008 A&A

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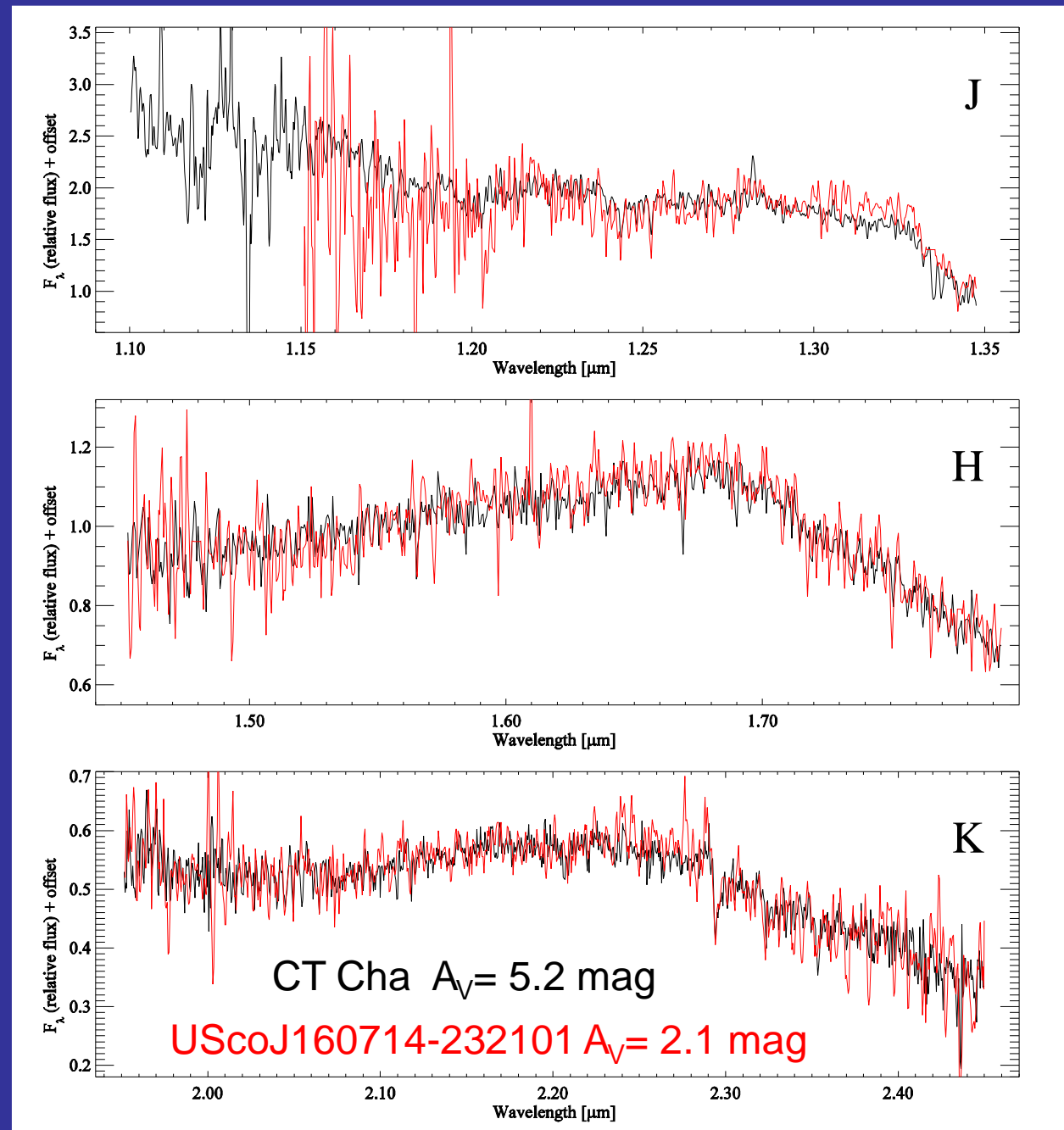
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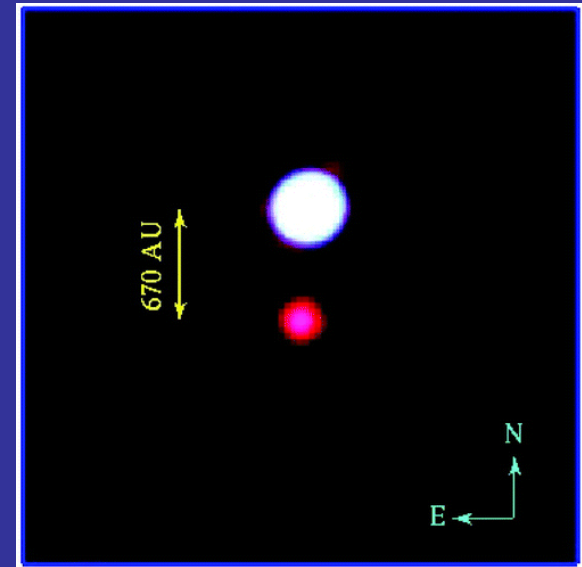
Lodieu et al. 2008



Schmidt, Neuhäuser, Seifahrt et al. 2008 A&A

UScoCTIO 108 b

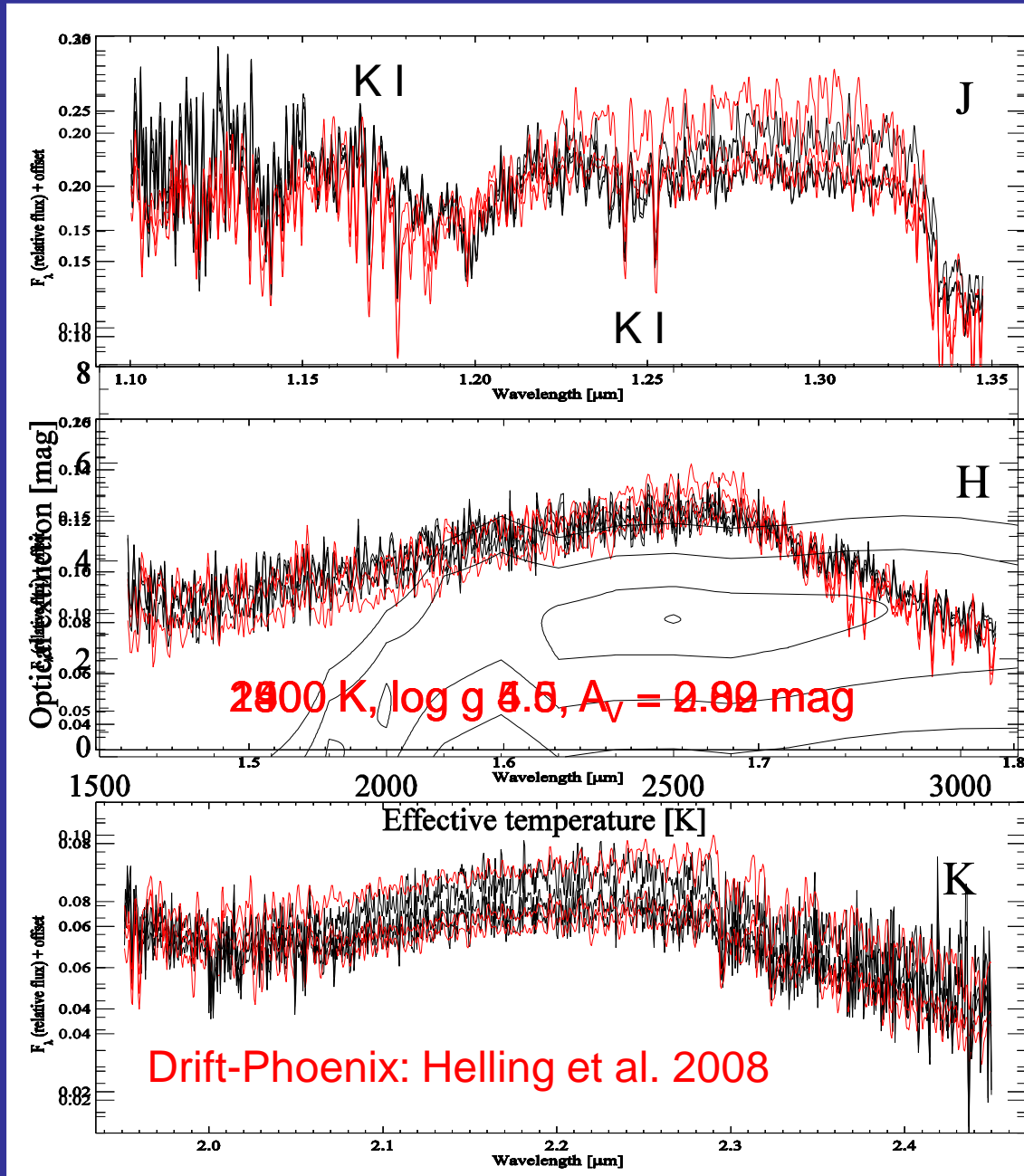
(Béjar et al. 2008)



and Drift-Phoenix:

$$T = 2500 \text{ K} \begin{matrix} + 400 \text{ K} \\ - 200 \text{ K} \end{matrix}$$

$$A_V = 2.8 \pm 0.8 \text{ mag}$$



Schmidt, Neuhäuser, Seifahrt 2011, in prep.

<u>Object</u>	<u>Temp.</u>	<u>A_v</u>
UScoJ155419-213543	2700 K	0.2 mag
UScoJ160648-223040	2700 K	0.2 mag
UScoJ160723-221102	2700 K	0.7 mag
UScoJ160847-223547	2700 K	0.7 mag
UScoJ154722-213914	2700 K	0.9 mag
UScoJ160830-233511	2700 K	1.6 mag
UScoJ161047-223949	2700 K	1.6 mag
UScoJ160603-221930	2650 K	2.0 mag
UScoJ160737-224247	2500 K	2.0 mag
UScoJ160818-223225	2500 K	2.0 mag
UScoJ160606-233513	2700 K	2.1 mag
UScoJ161302-212428	2500 K	2.1 mag
UScoJ160714-232101	2600 K	2.3 mag
UScoJ160727-223904	2500 K	2.5 mag
UScoJ160828-231510	2700 K	2.6 mag
UScoCTIO 108 b	2500 K	2.8 mag
UScoJ161441-235105	2600 K	2.9 mag
UScoJ163919-253409	2600 K	4.2 mag
UScoJ161228-215936	2700 K	4.3 mag
UScoJ160918-222923	2500 K	4.6 mag

T = 2350 + 100 K
- 400 K

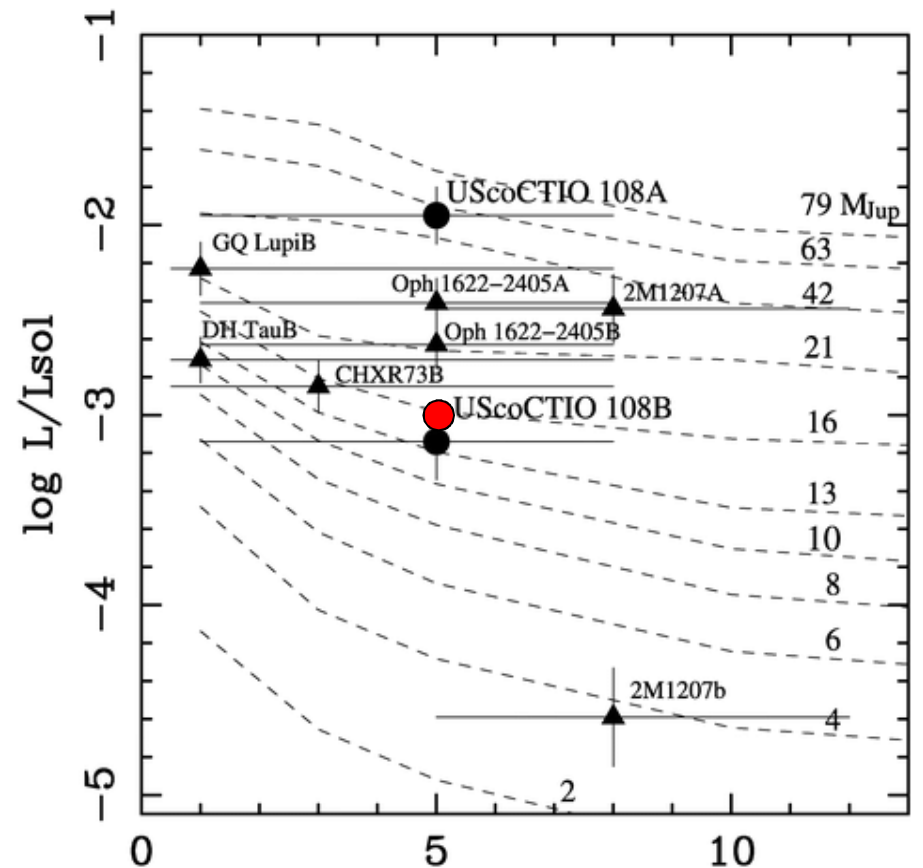
→

T = 2500 + 400 K
- 200 K

log(L/L_⊙) = -3.14 ± 0.20

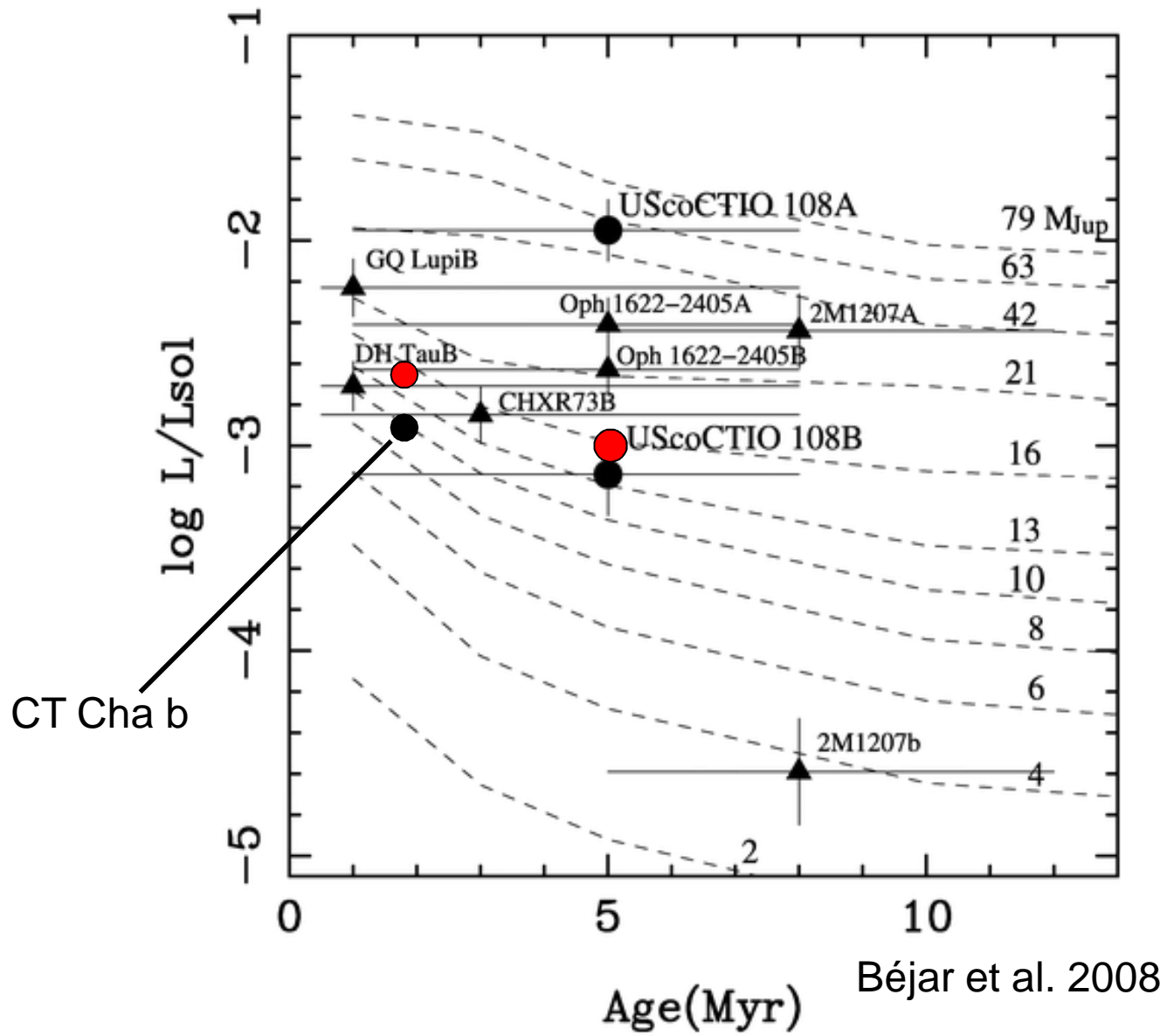
→

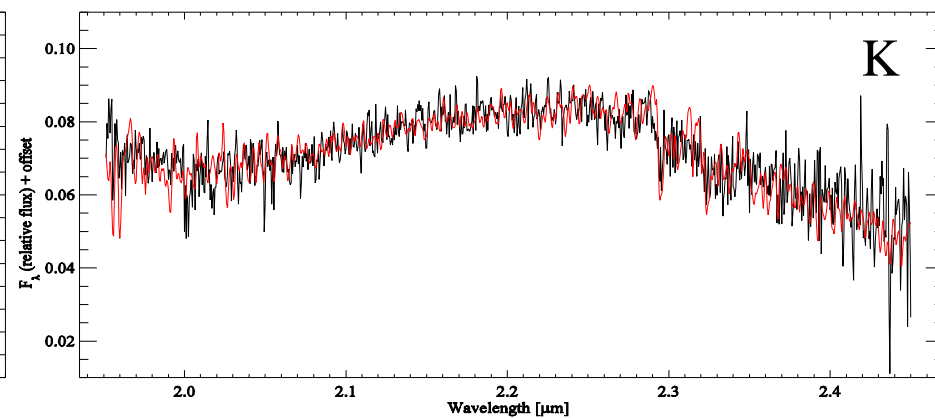
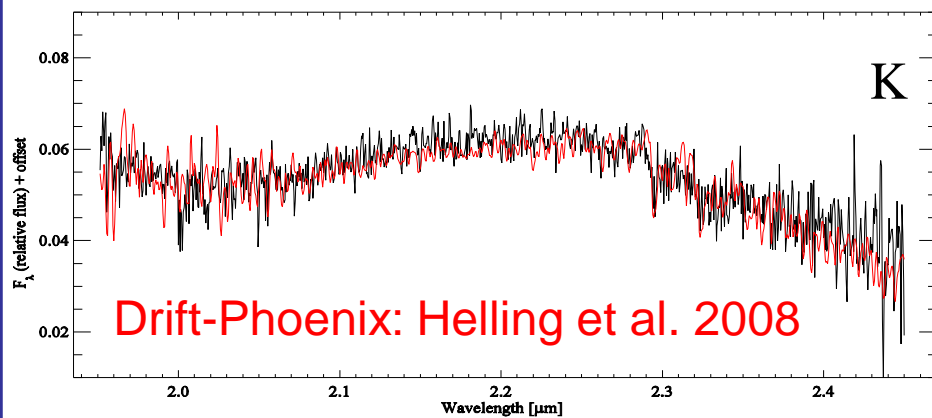
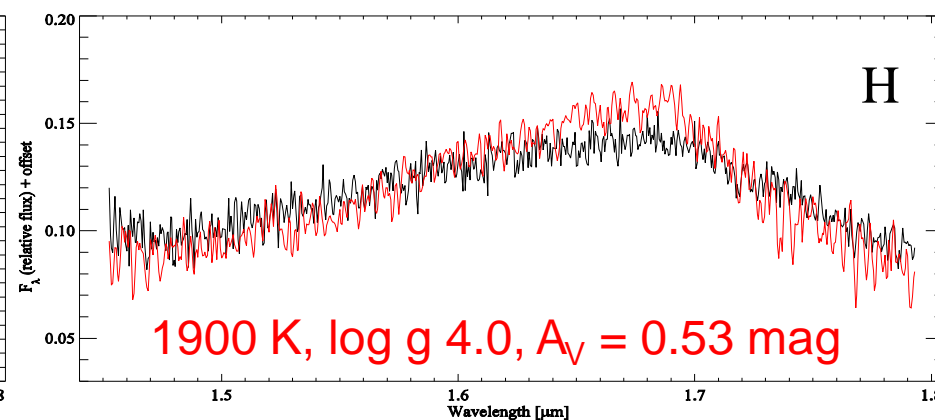
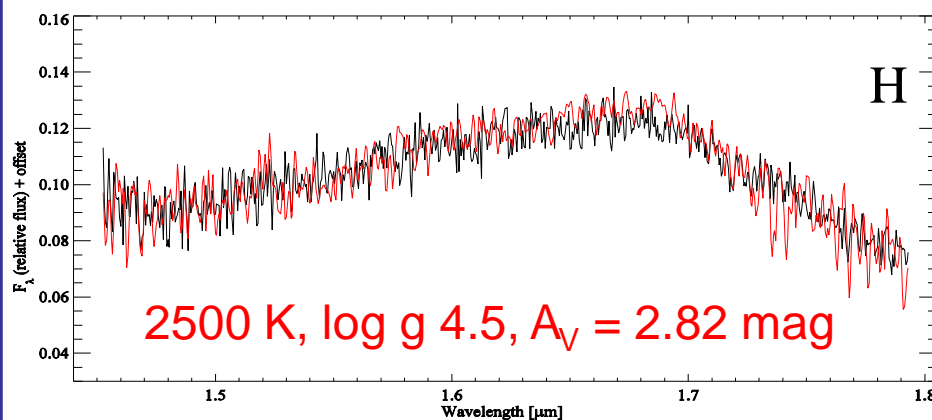
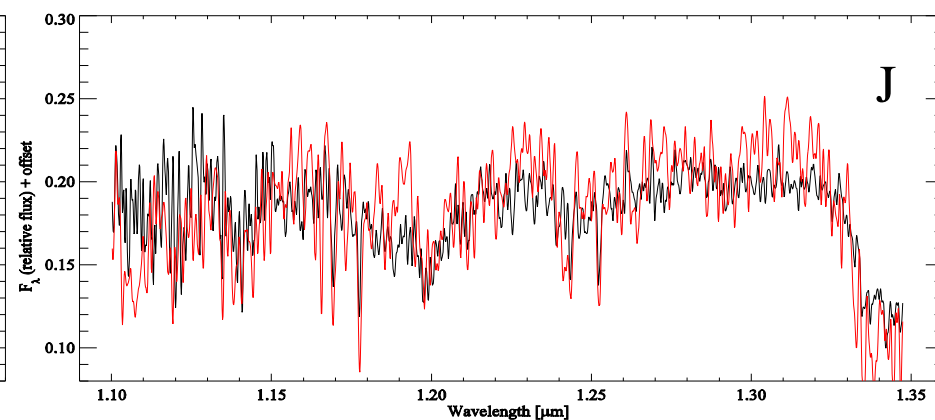
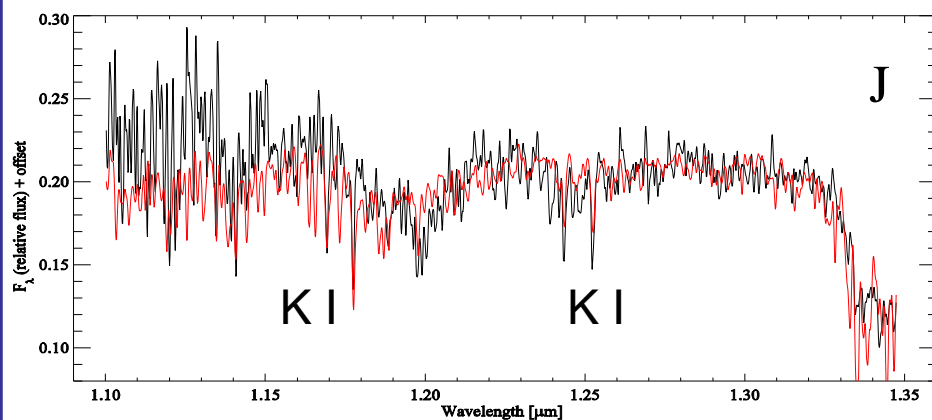
-3.01 ± 0.25



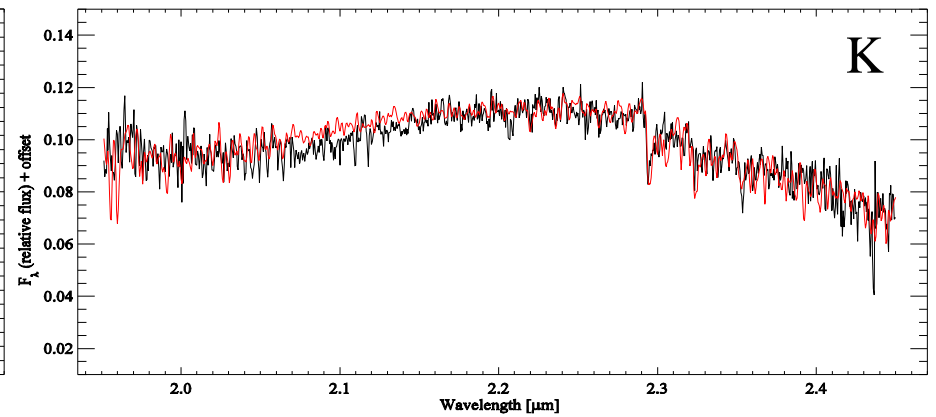
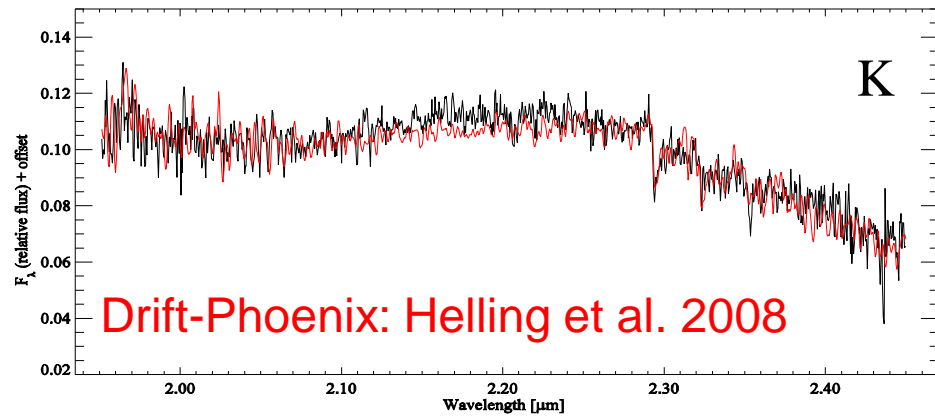
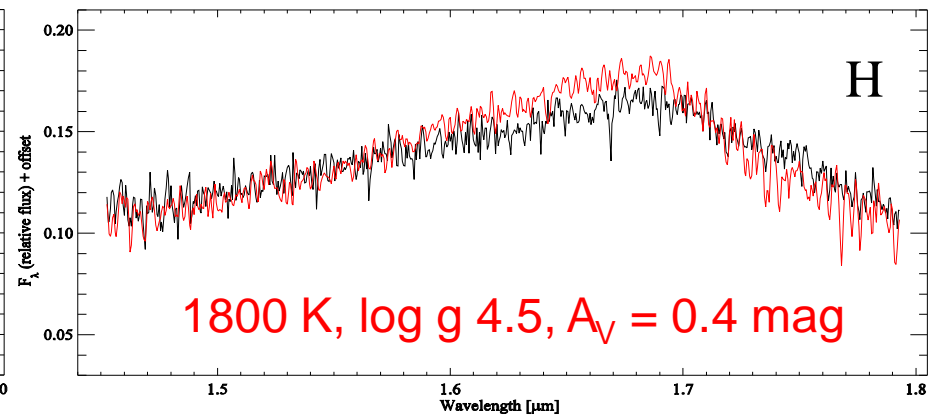
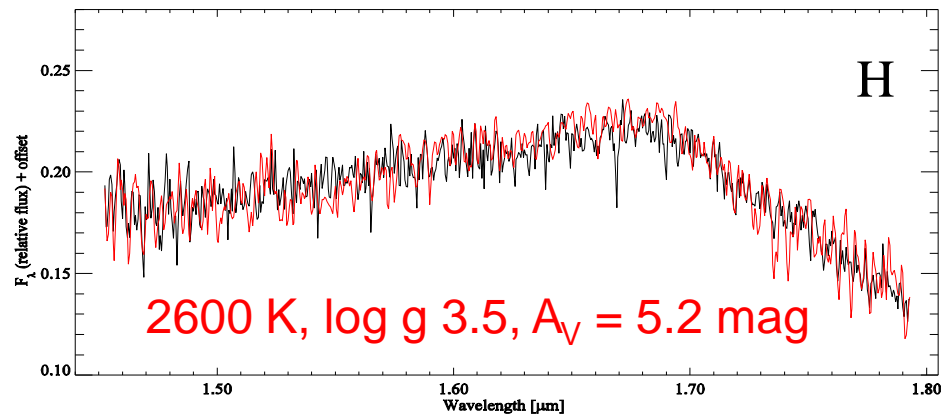
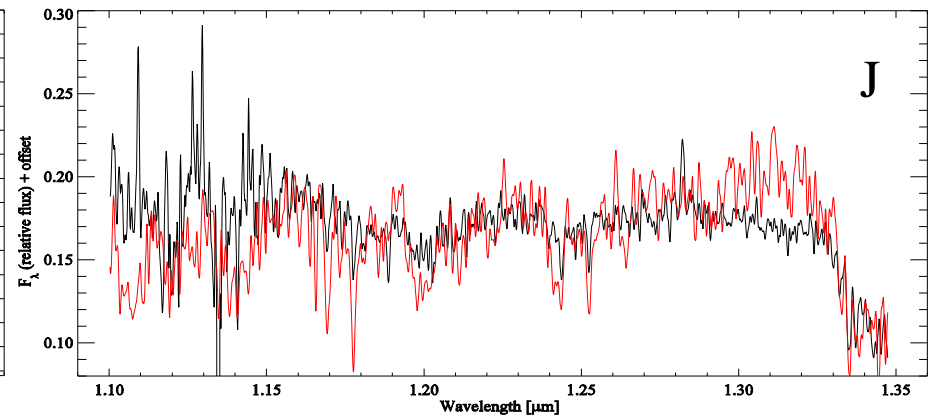
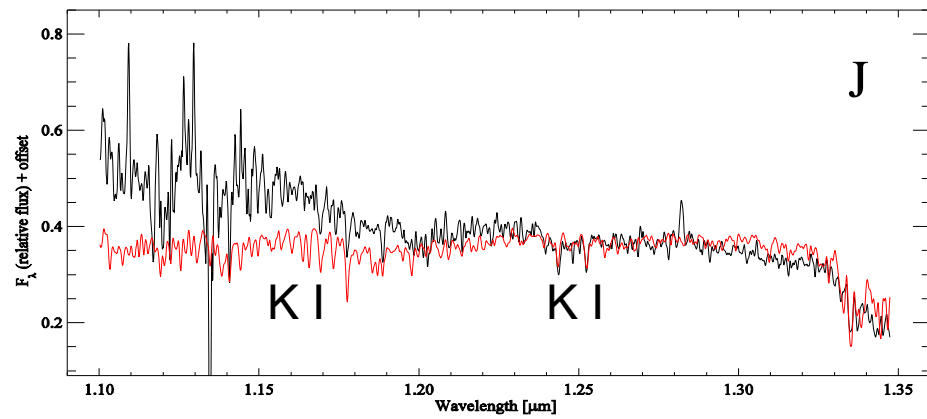
Béjar et al. 2008
(Baraffe model)

Lodieu et al. 2008 **Béjar et al. 2008 (new fit)**



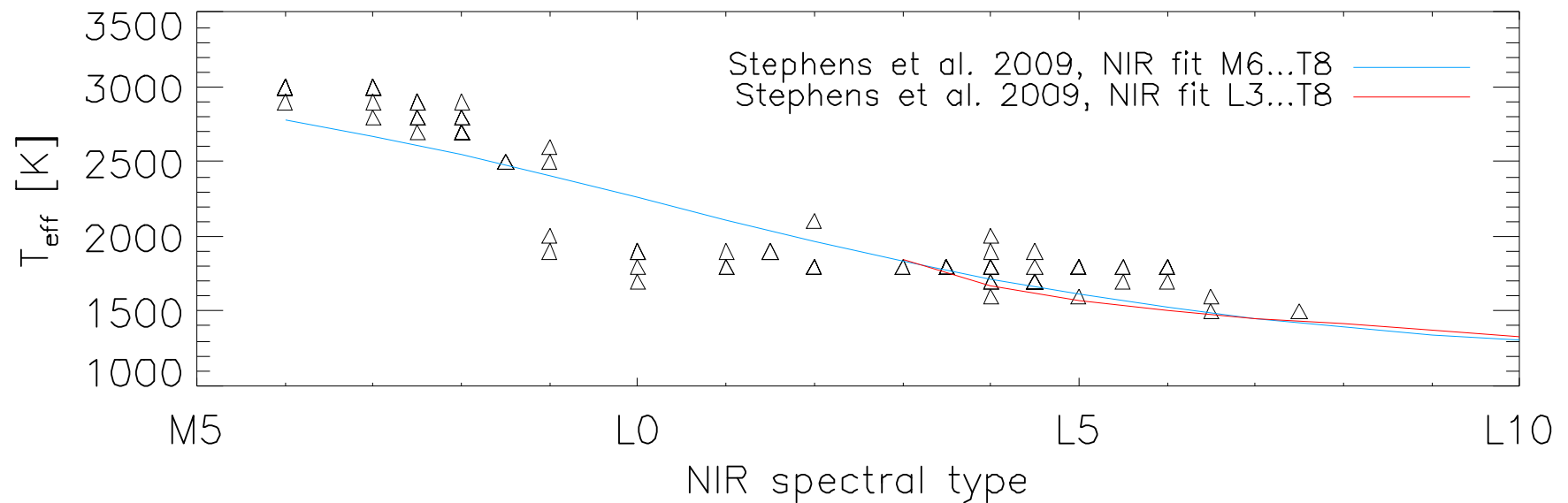
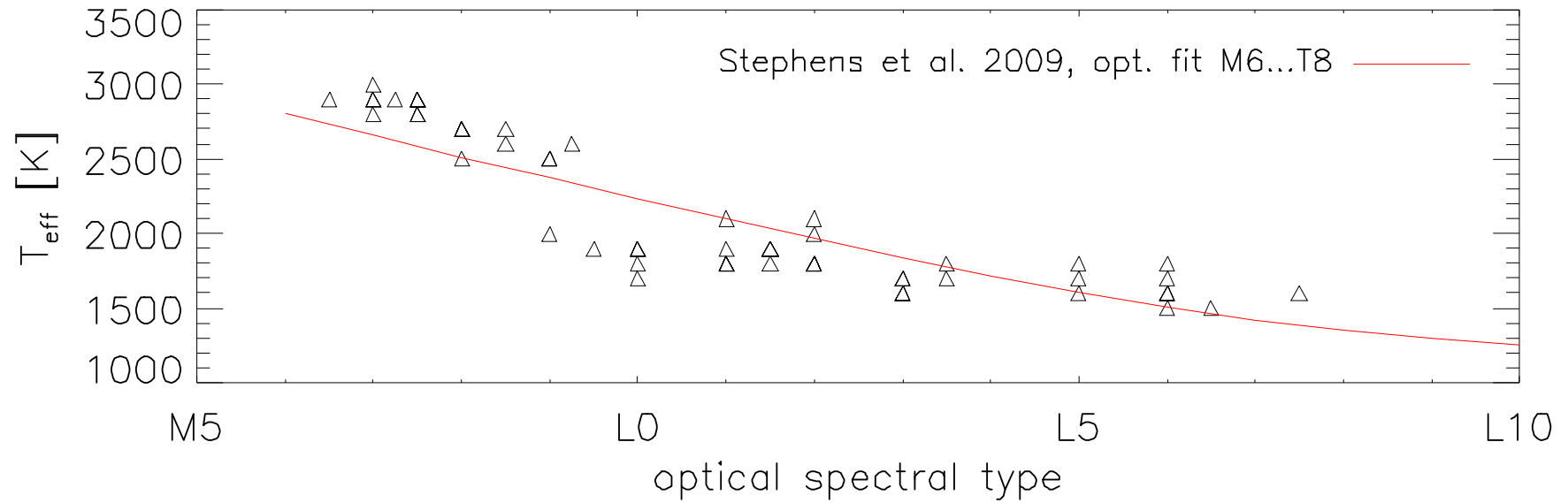


Schmidt, Neuhauser, Seifahrt 2012, in prep.



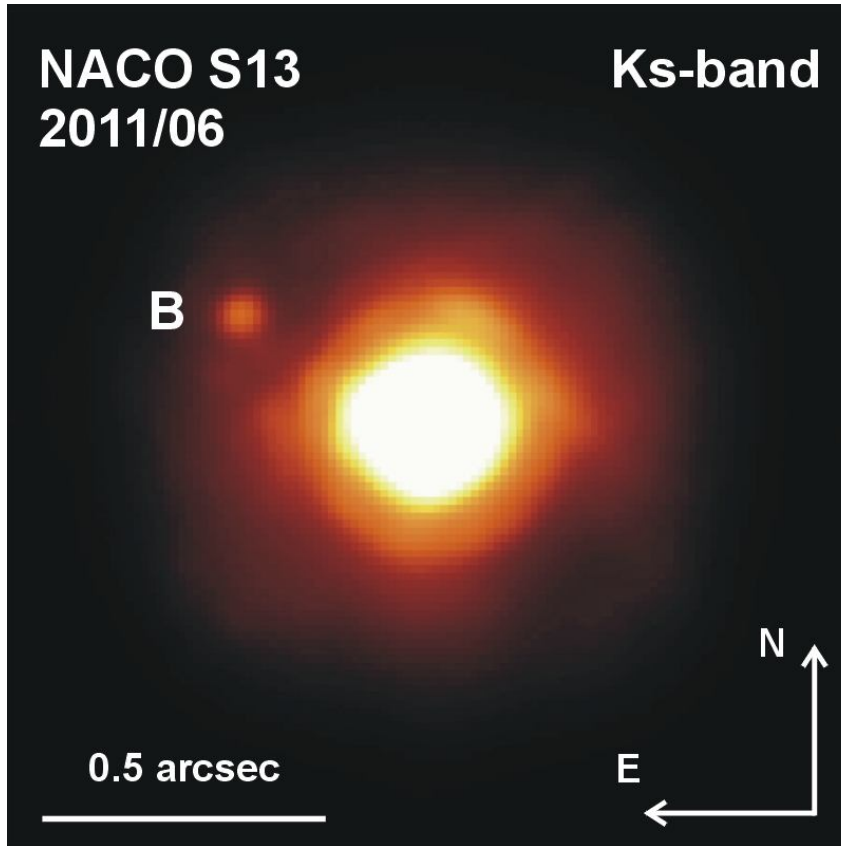
Schmidt, Neuhauser, Seifahrt 2012, in prep.

Witte et al. 2011 – Comparison with the 108 suited (e.g. single) BDs from DwarfArchives

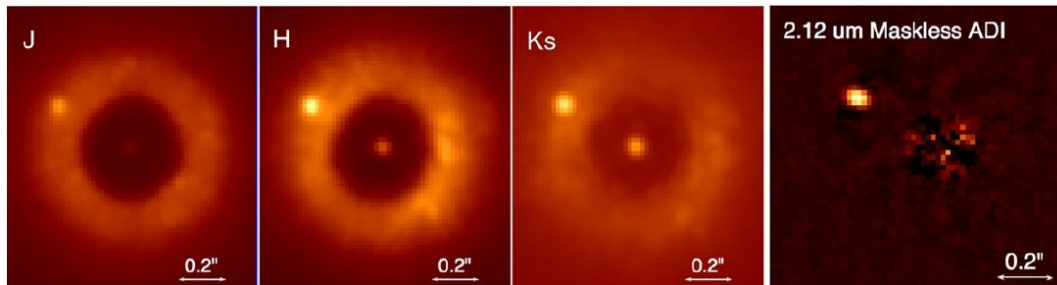
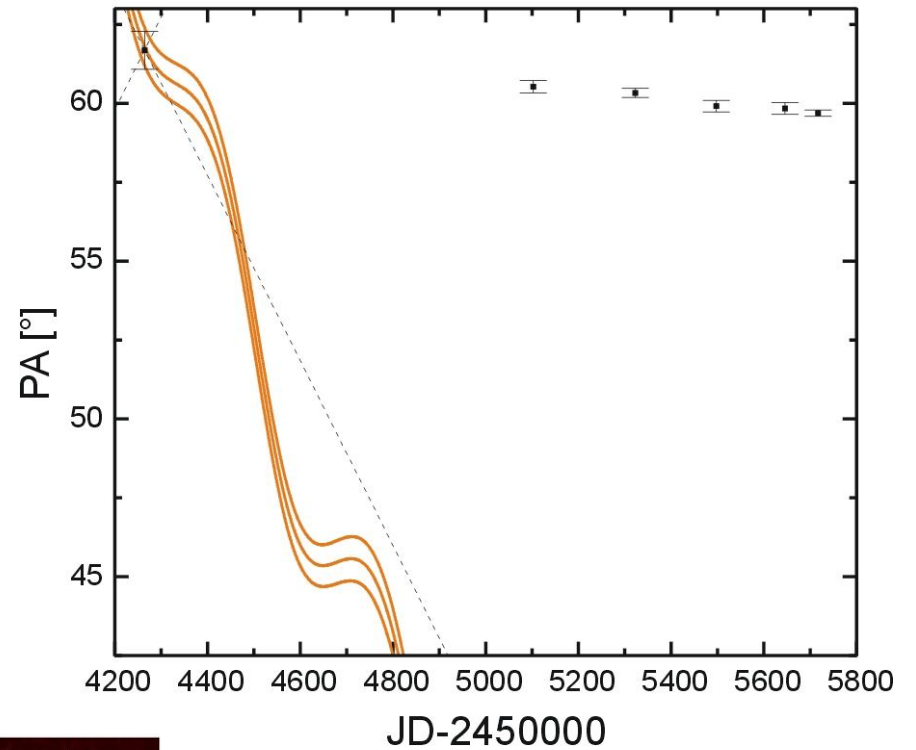


PZ Tel

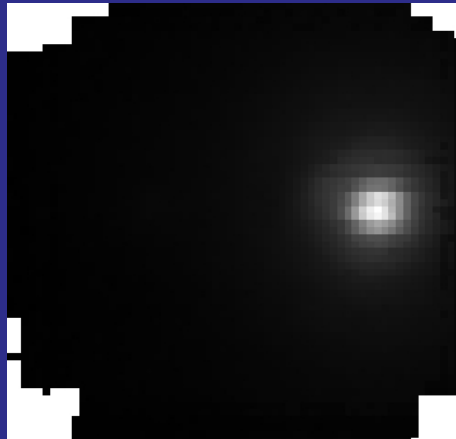
12 Myrs (in β Pic moving group)



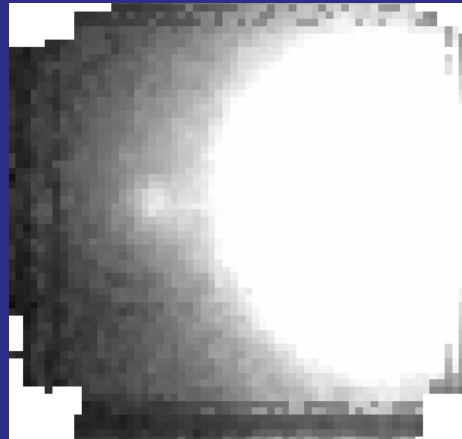
New astrometric data \rightarrow please see
Poster S2/P15



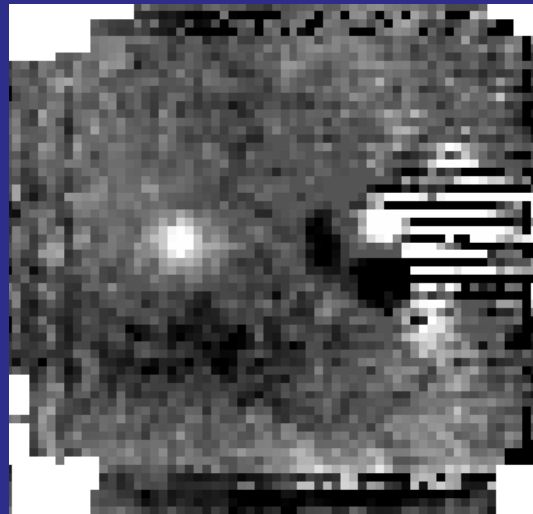
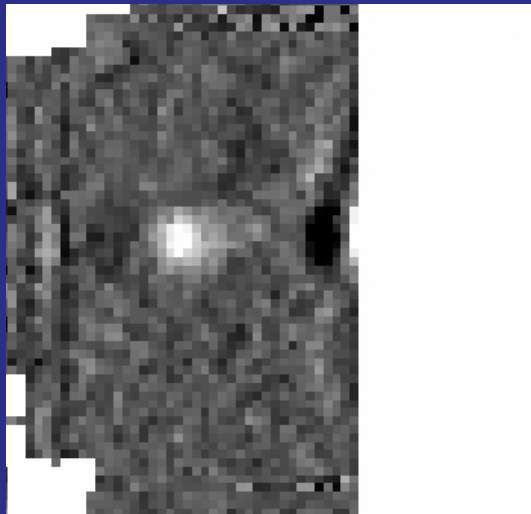
Simultaneously found by
Mugrauer et al. 2010 &
Biller et al. 2010



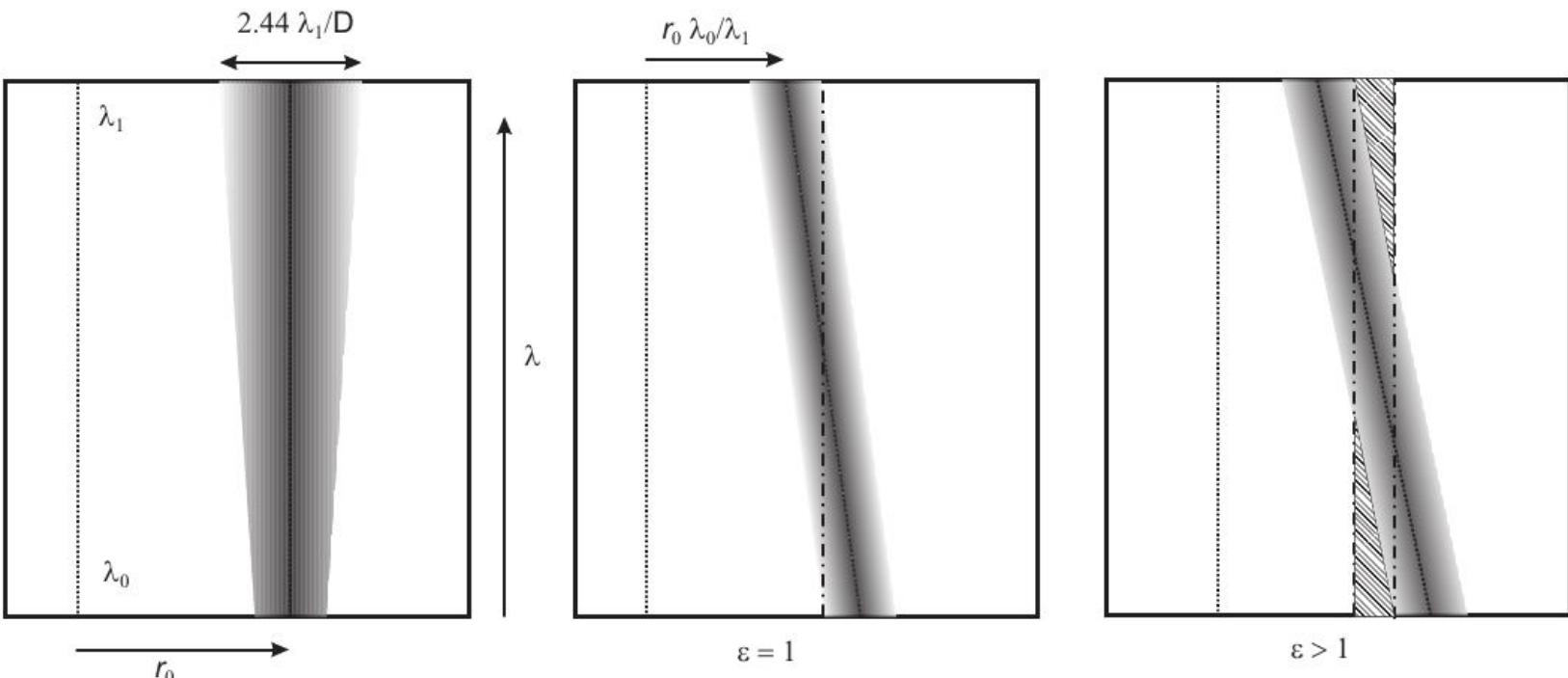
Polynomial removal



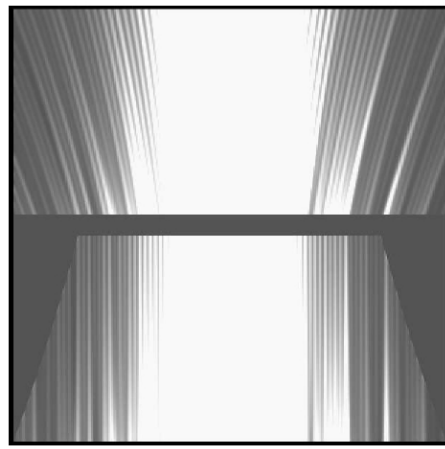
radial symmetry removal



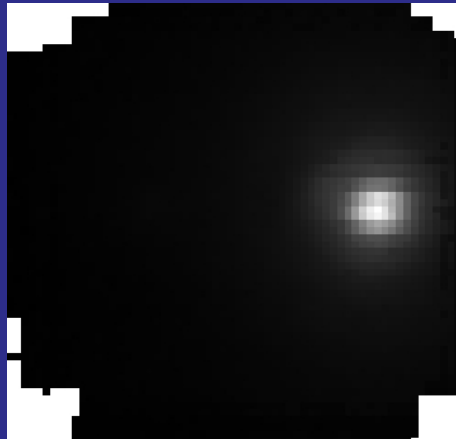
Spectral deconvolution



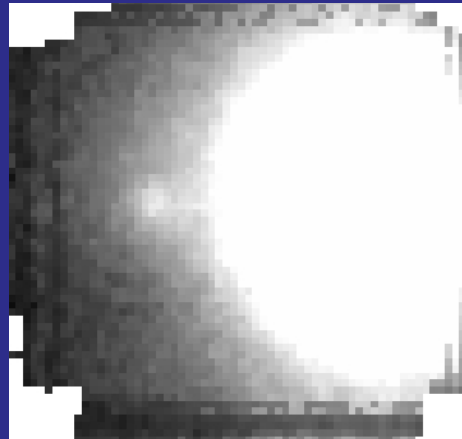
Thatte et al. 2007



Based on Sparks & Ford 2002



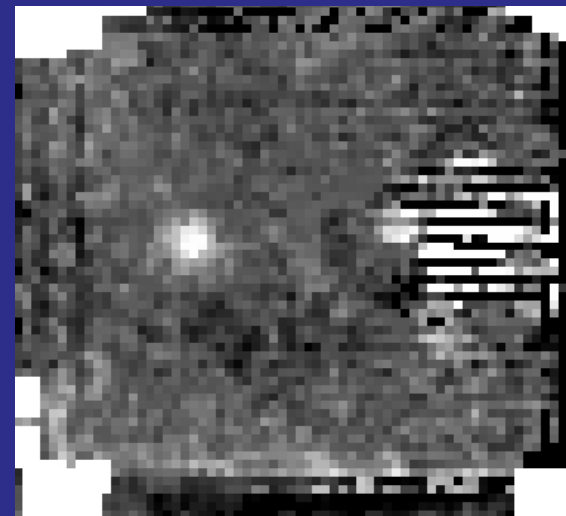
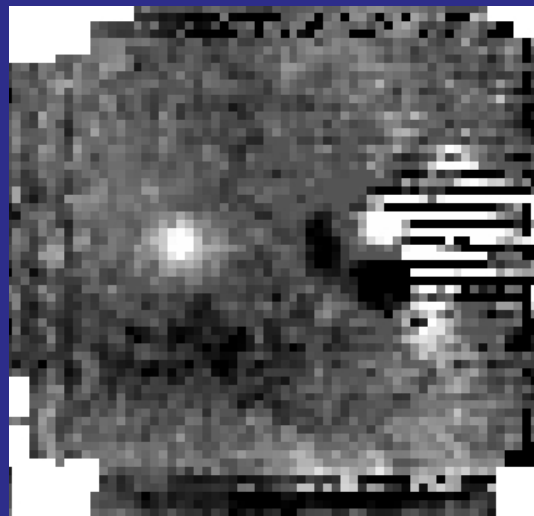
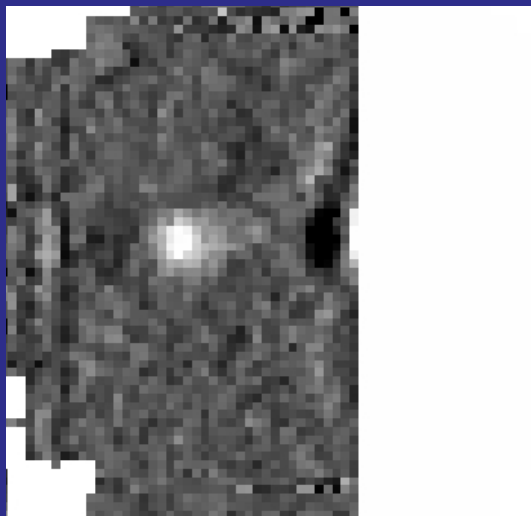
Polynomial removal

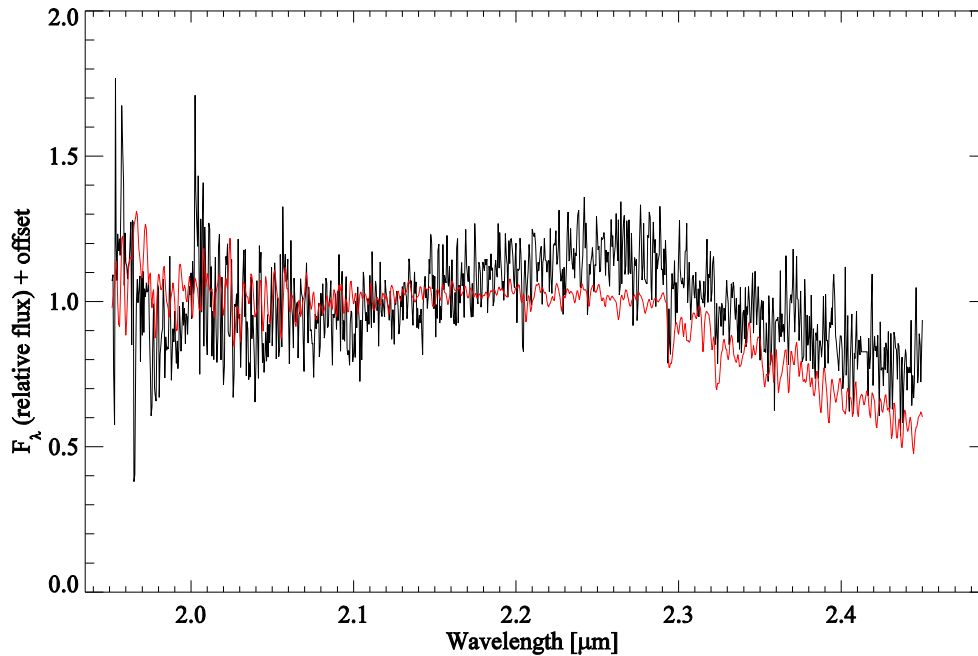


radial symmetry removal



spectral deconvolution

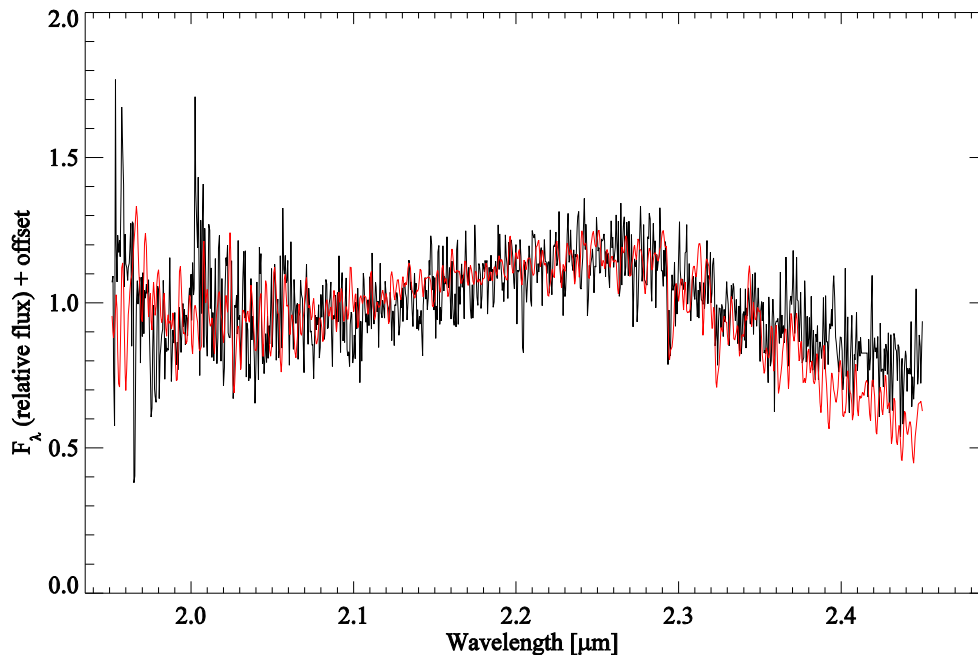




Values from Jenkins et al. 2012:

age lithium: 7 Myr
 age chromosph. activity: 26 Myr
 age evol. models: 22 Myr

3000 K, $\log g$ 4.5, $[Fe/H] = 0.0$



Best fitting values close to values
 of Mugrauer et al. 2010 &
 Biller et al. 2010:

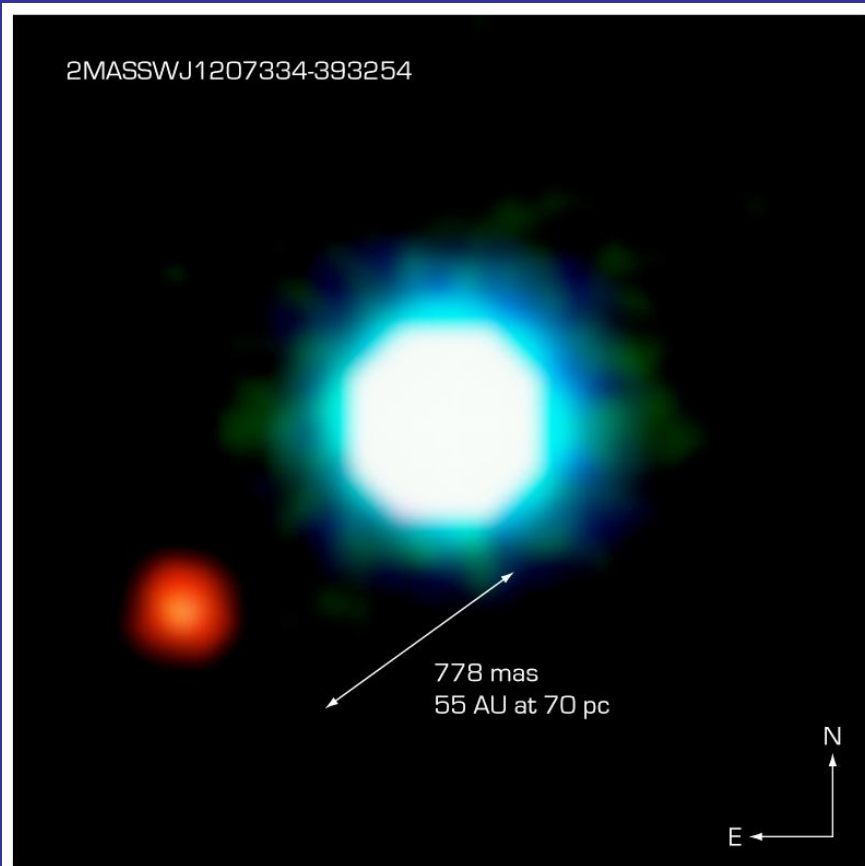
2600 K, $\log g$ 3.5, $[Fe/H] = 0.3$

consistent with 30 Mjup object
 at 12 Myr

Schmidt, Mugrauer, 2012, in prep.

Chauvin et al. (2004): Planetary companion to 2M1207

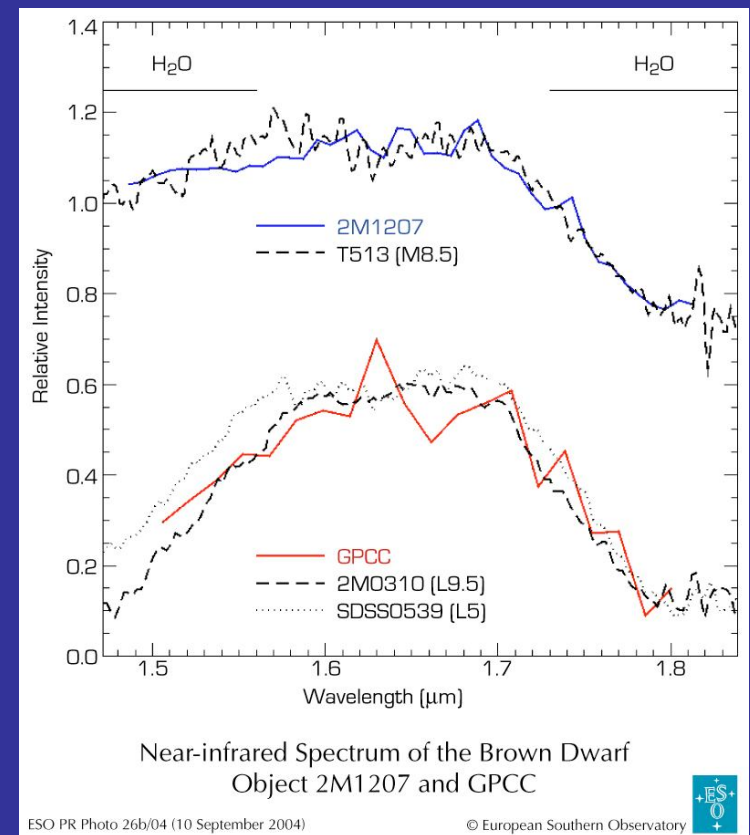
2M1207: Possible brown dwarf in TWA, 5 to 12 Myrs, 50 to 100 pc



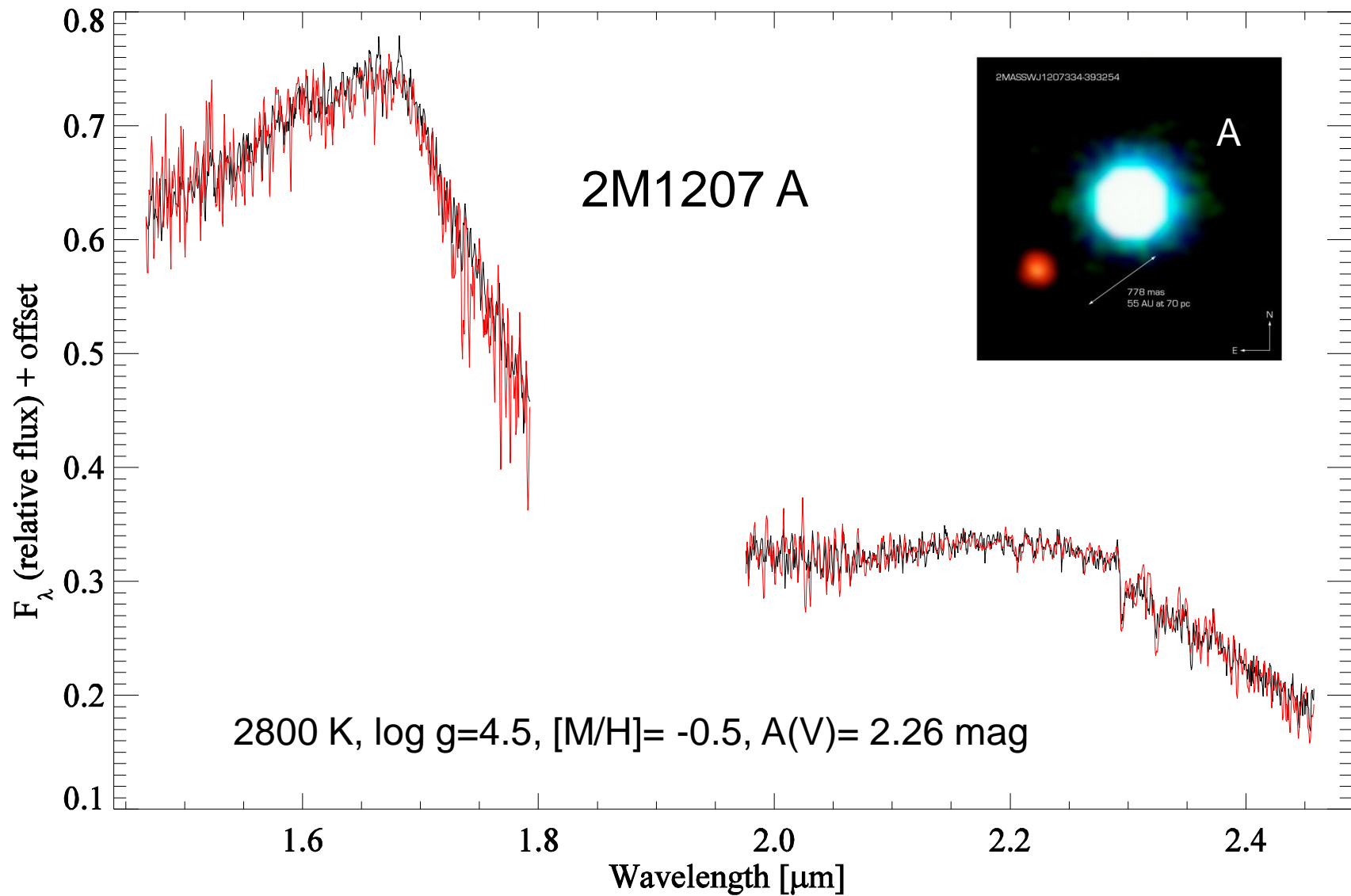
NACO Image of the Brown Dwarf Object 2M1207 and GPCC

ESO PR Photo 26a/04 (10 September 2004)

© European Southern Observatory

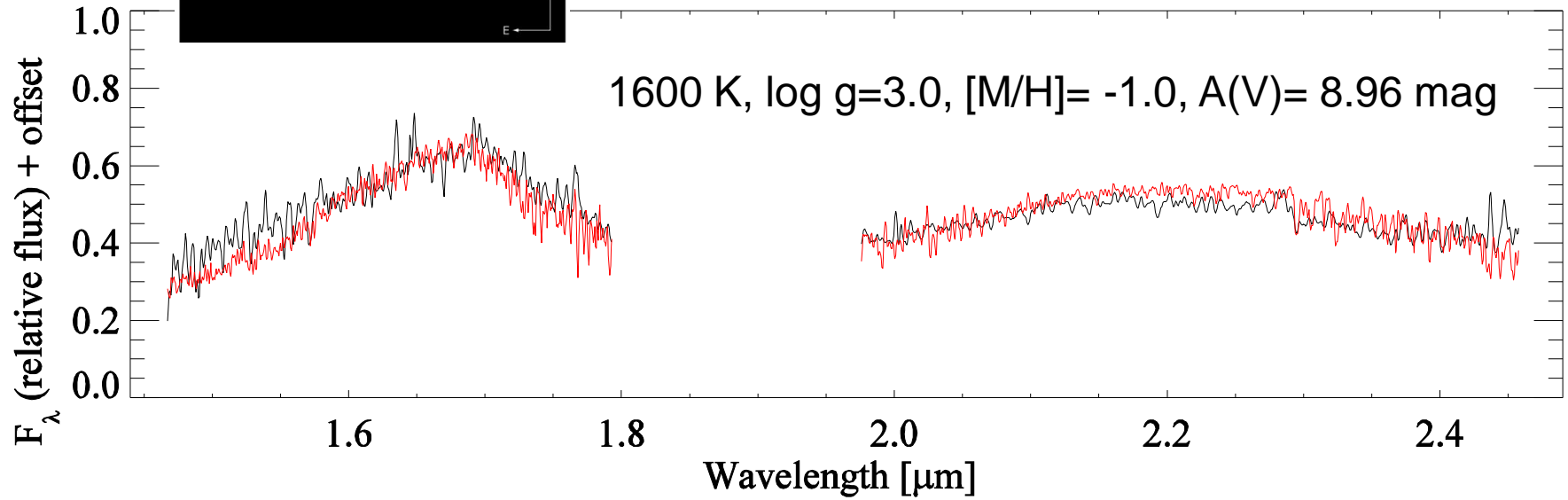
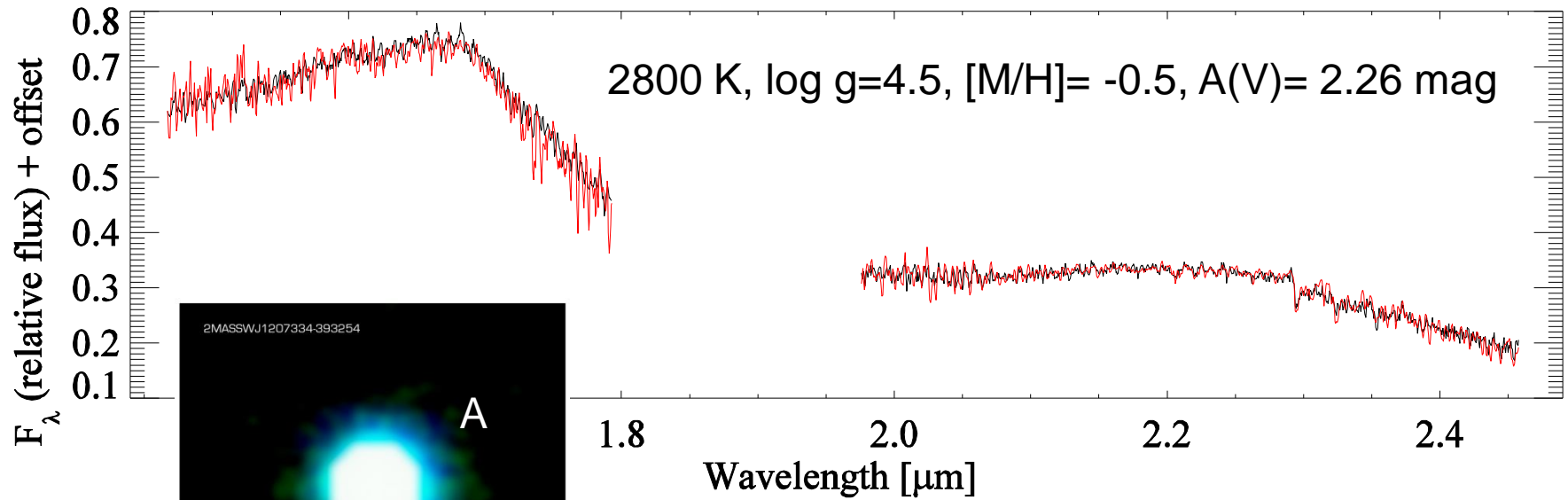


Common proper motion (Chauvin et al. 2005)

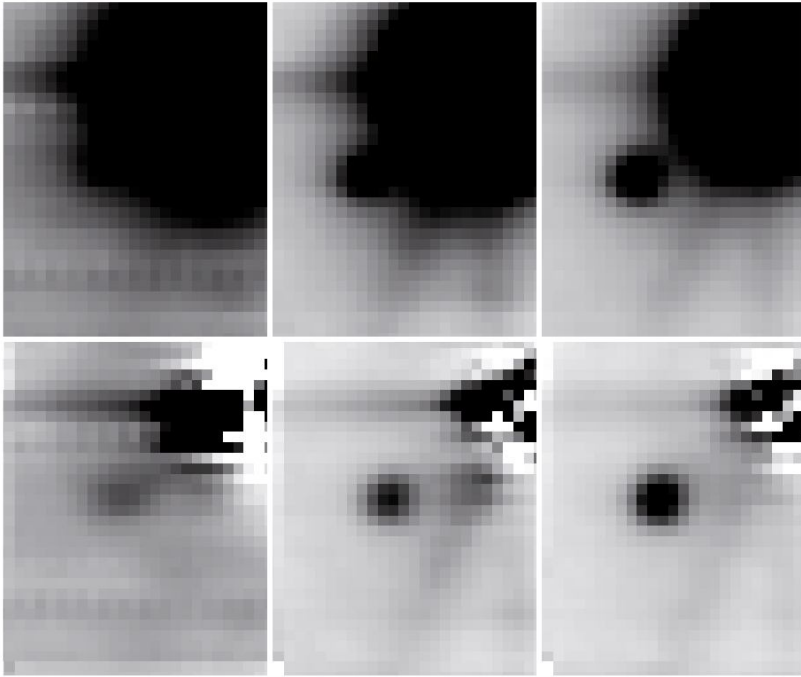


SINFONI spectrum from ESO archive

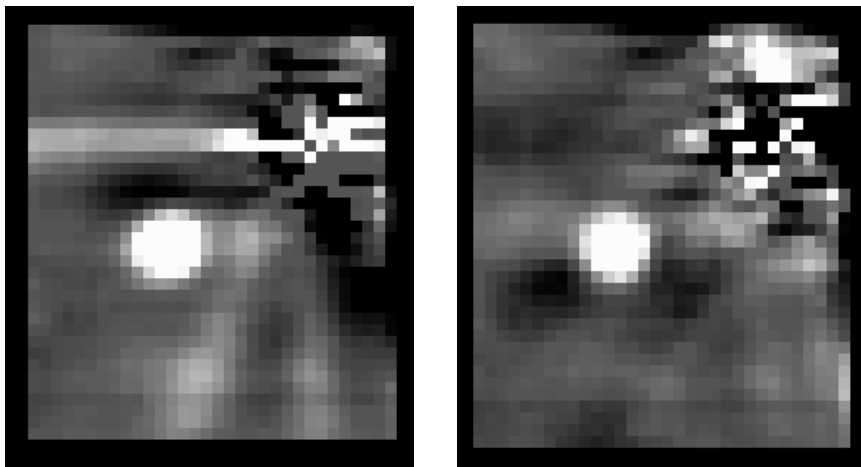
Drift-Phoenix: Helling et al. 2008



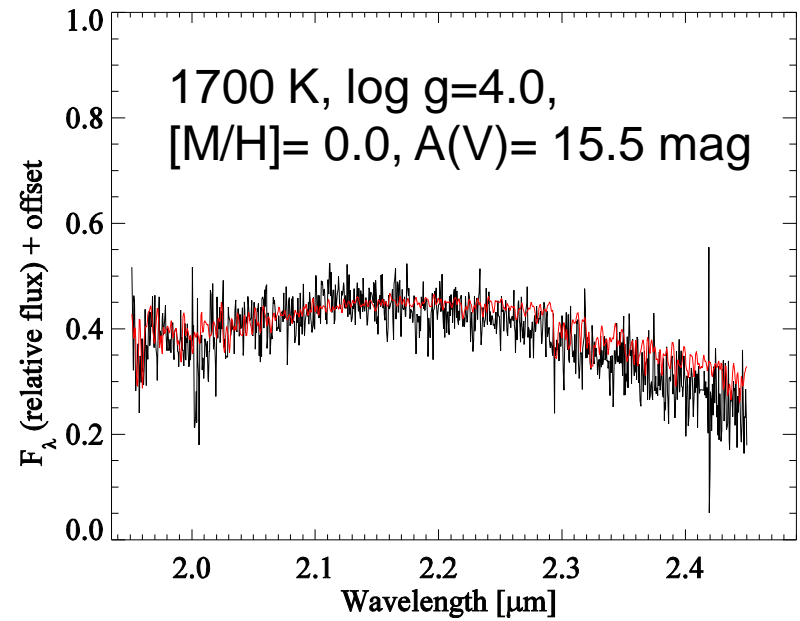
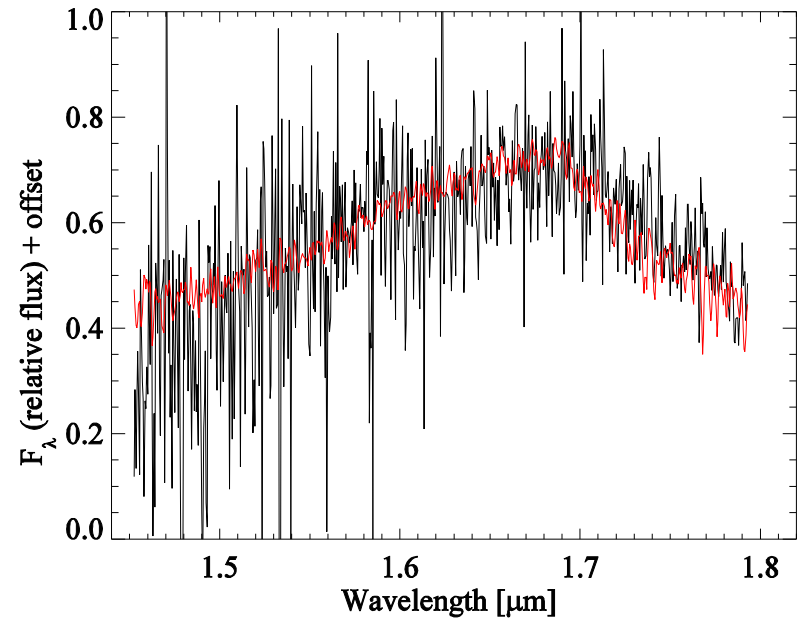
Patience et al. 2010



spectral deconvolution



Schmidt, Quanz, Seifahrt, 2012, in prep.



2M1207 b: $A_V = 8.96$ mag

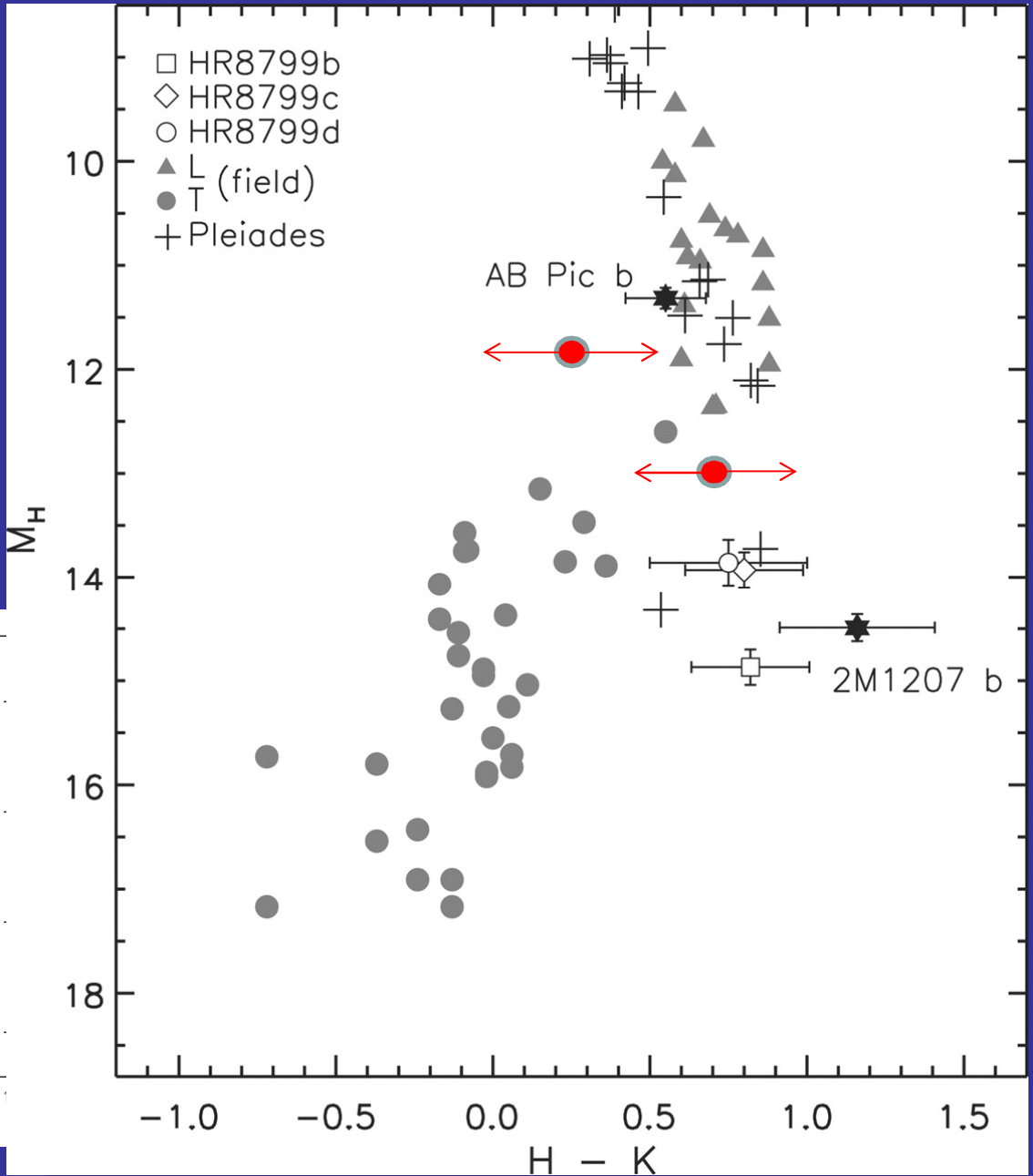
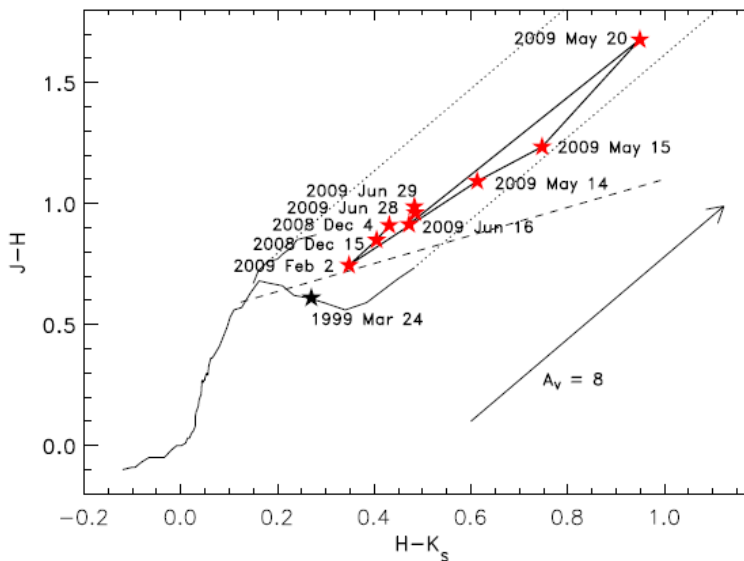
$\log(L/L_\odot) = -4.54 \pm 0.1$
(Mamajek et al. 2007)

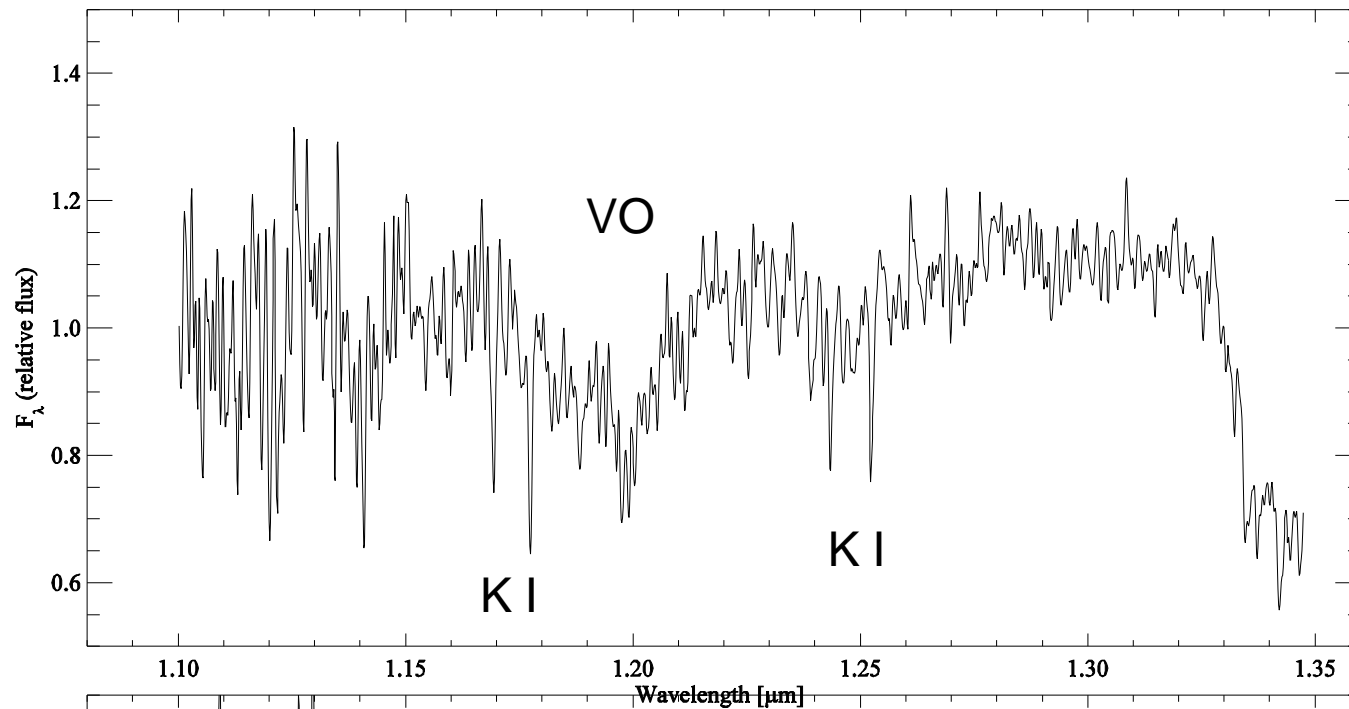
→ -4.14 ± 0.2

Underluminosity:

4 – 7 times @ 5 – 10 Myr

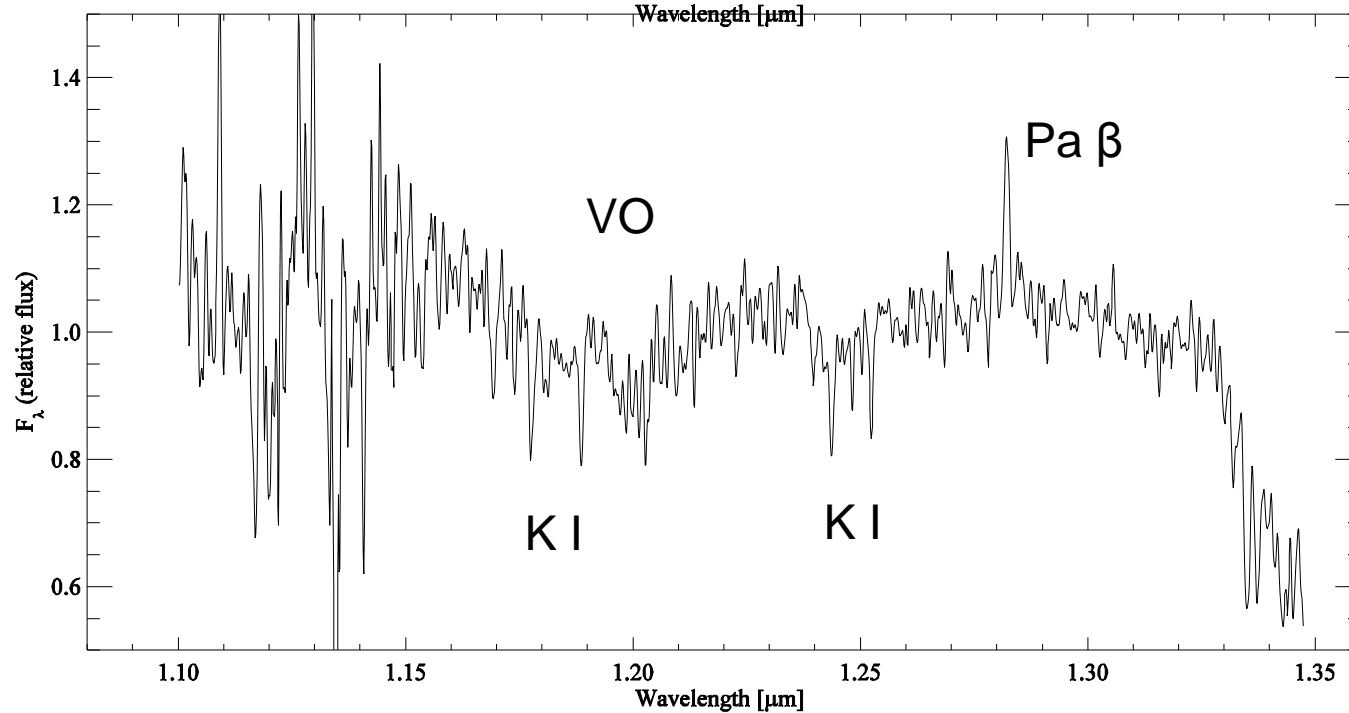
→ 1 – 4 times @ 5 – 10 Myr





UScoCTIO 108 b

~ 5 Myr



CT Cha b

~ 2 Myr



Thank you
for your attention!!

Acknowledgements

- Nikolaus Vogt, Ana Bedalov
- David Lafrenière, Nicolas Lodieu, J. Davy Kirkpatrick, Kevin Luhman
- Christiane Helling, Soeren Witte, Peter Hauschildt, Takashi Tsuji, Mark Marley
- Evangelisches Studienwerk e.V. Villigst
- FONDECYT
- DFG, DAAD

